

Flare-related radio emission: a kinetic point of view

Carine Briand

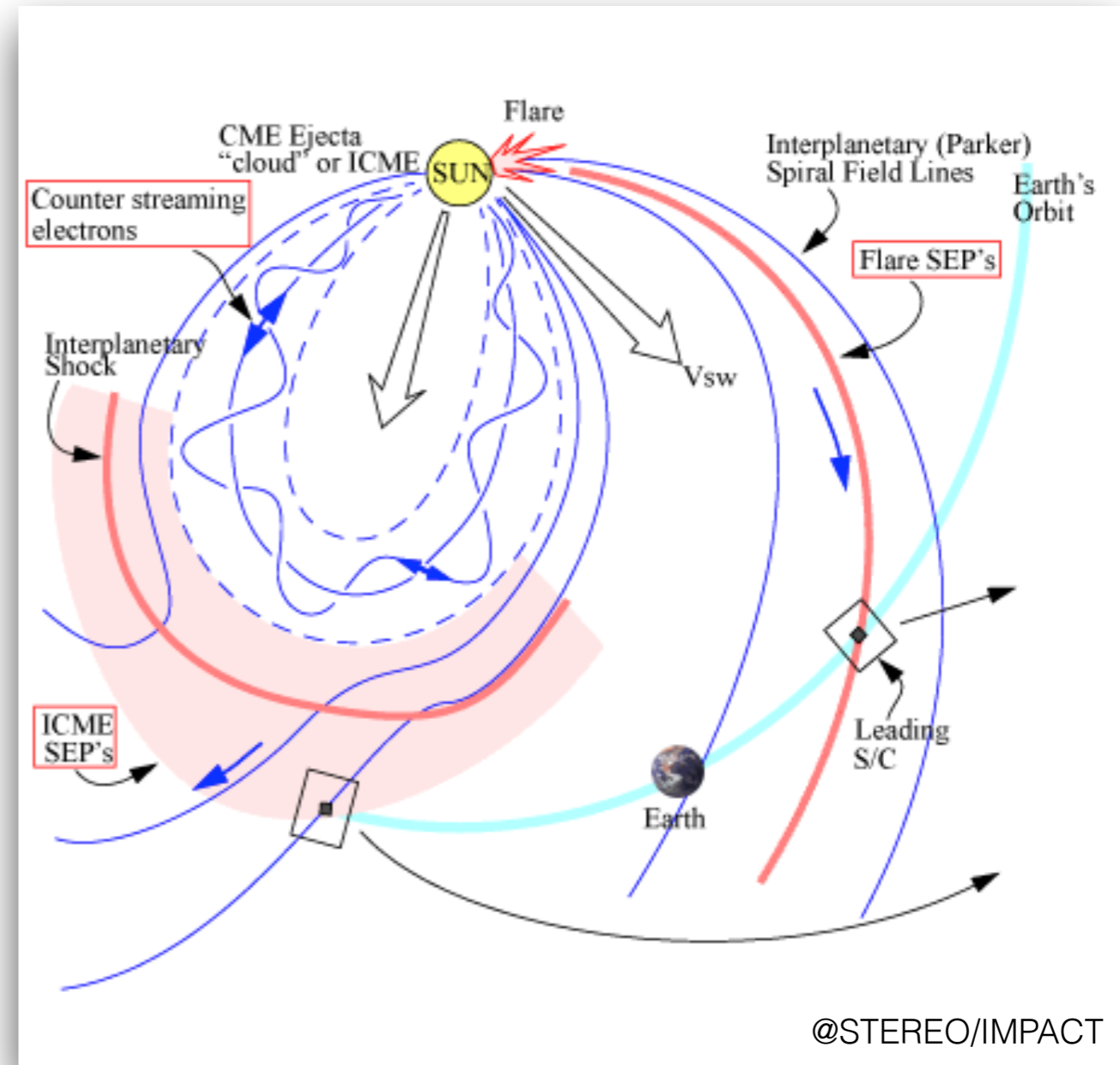
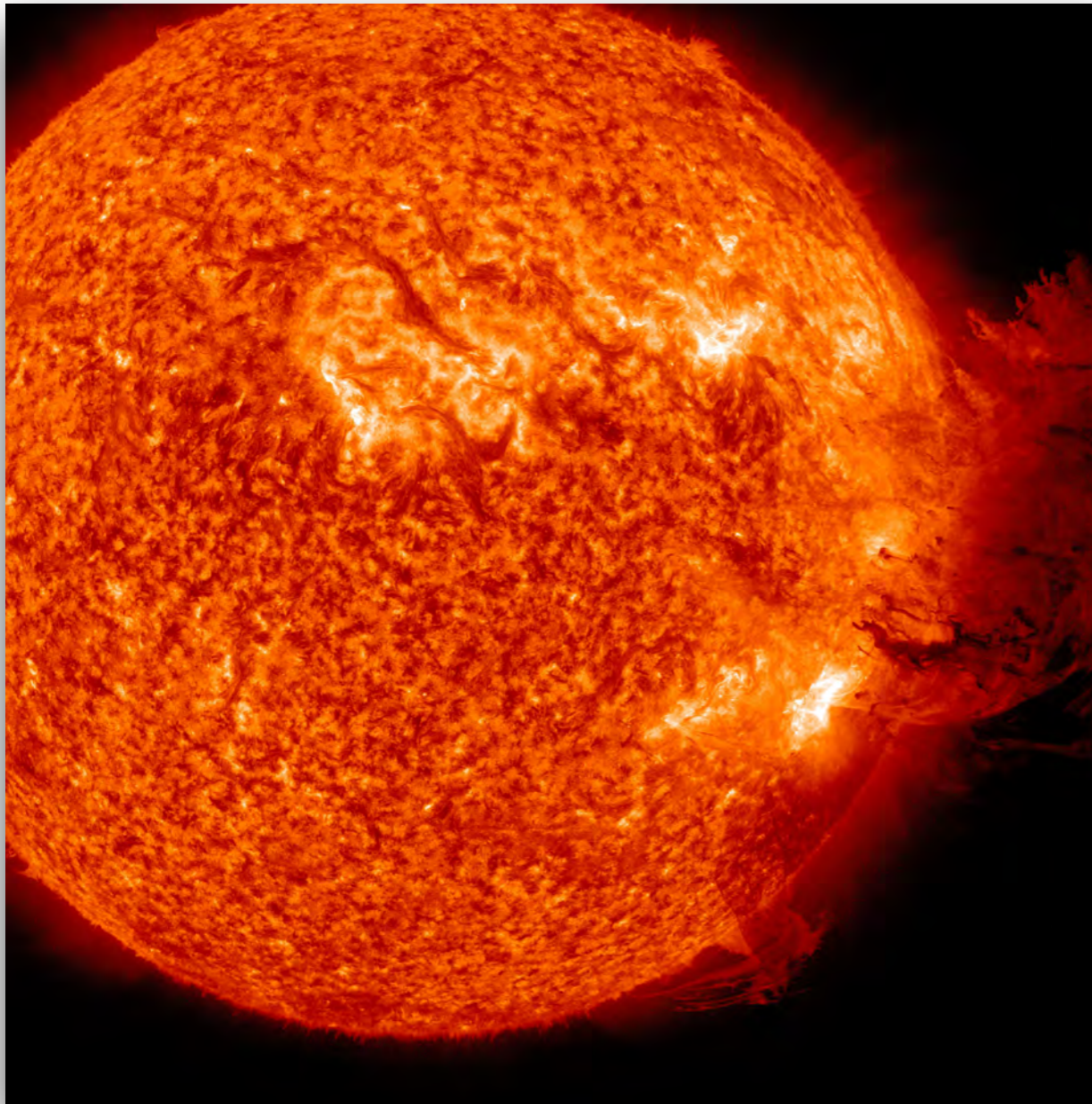
Paris Observatory, LESIA

& Co-workers

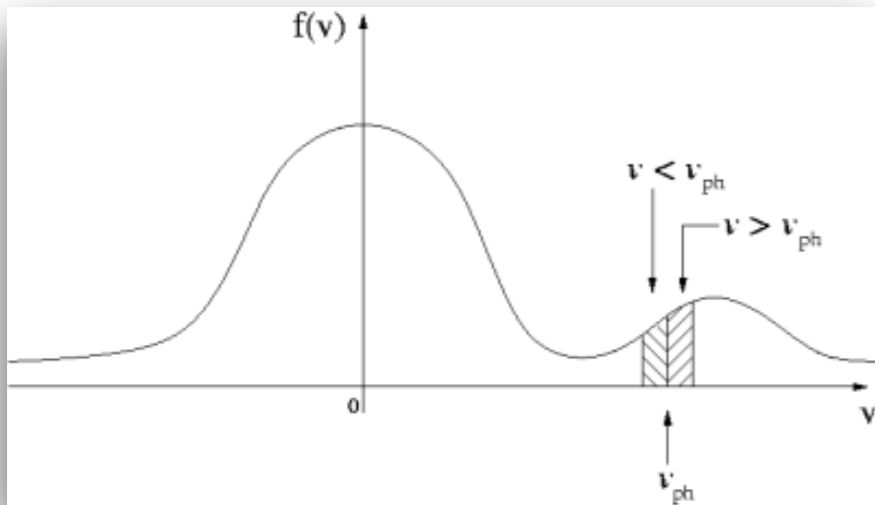
Pierre HENRI, LPC2E, France

Francesco Califano, Pisa Univ., Italy

Eruptive Sun



Fundamental of beam plasma interaction



Suprathermal electrons: free energy for plasma instability
(bump-on-tail)

Growth of Langmuir waves
(local plasma frequency - plasma density)

Non Linear Saturation and generation of EM emissions
(resonant and parametric (threshold) processes)

Electromagnetic Decay

$$L \rightarrow T_{fp} + LF$$

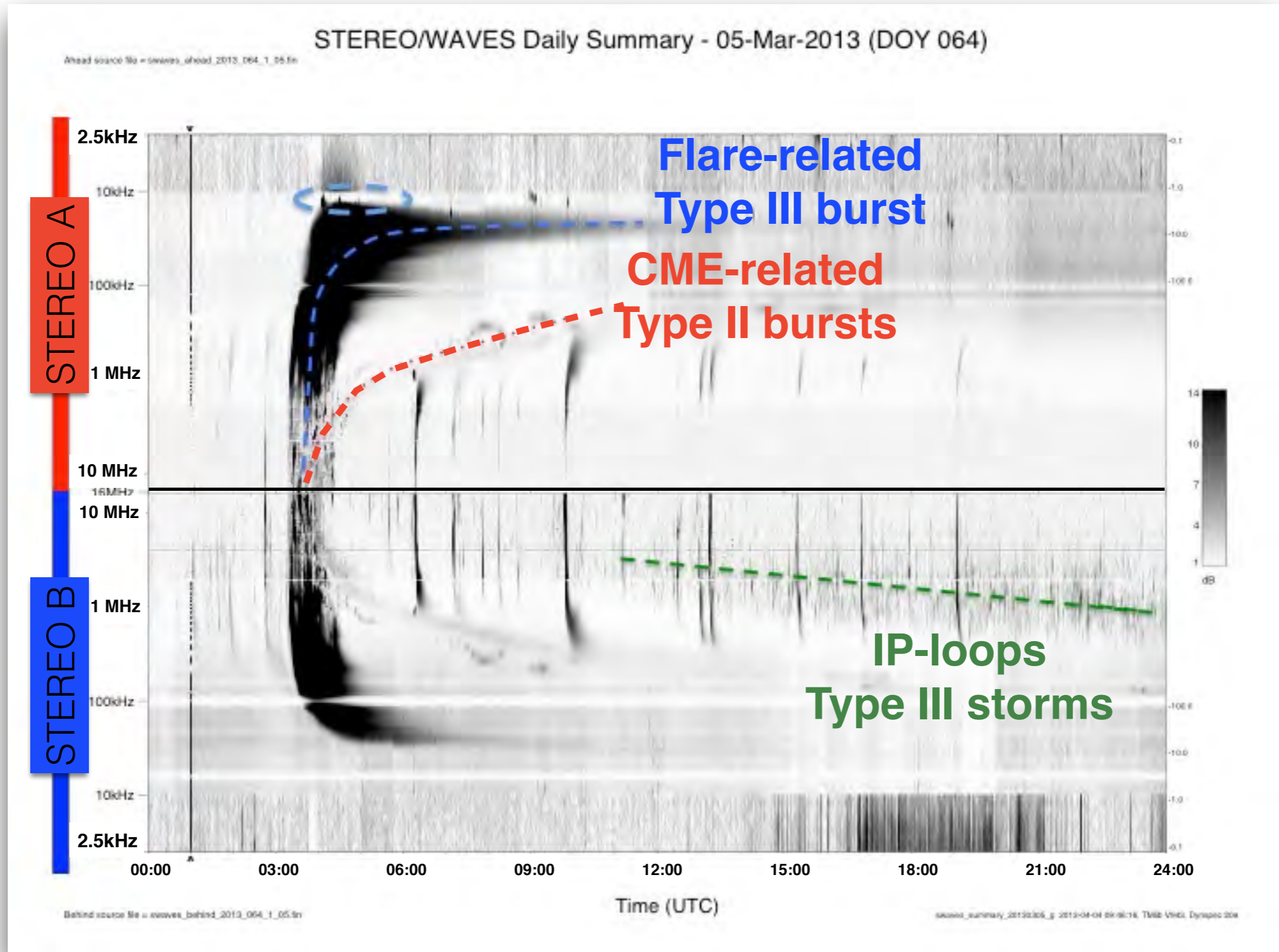
Electrostatic Decay & coupling

$$L \rightarrow L' + LF$$

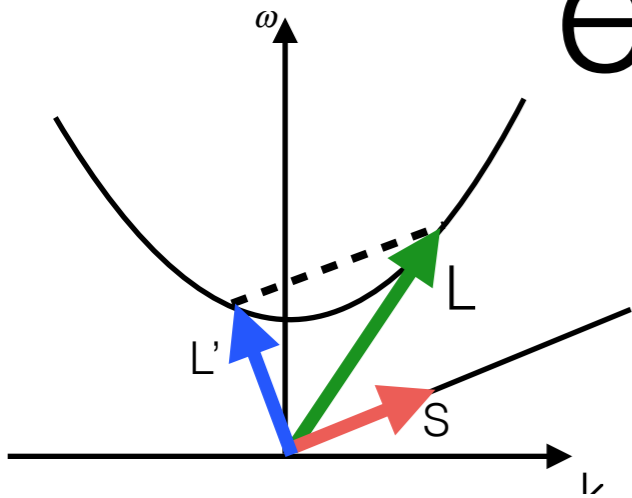
$$L' + L \rightarrow T_{2fp}$$

Lowest threshold

Solar radio emissions



Fundamental of Langmuir electrostatic decay



$$L \rightarrow L' + LF$$

$$L' + L \rightarrow T_{2fp}$$

LF = Ion Acoustic Waves

Resonant interaction:

- (i) conservation of momentum: $\mathbf{k}_L = \mathbf{k}_{L'} + \mathbf{k}_S$
- (ii) conservation of energy: $\omega_L = \omega_{L'} + \omega_S$
- (iii) phase locking: $\varphi_L = \varphi_{L'} + \varphi_S$

Threshold:

$$\gamma_{LED} > \sqrt{\gamma_{L'} \gamma_S}$$

$$\frac{\epsilon_0 E^2}{nk_B T} > 8 \frac{\gamma_{L'}}{\omega_{L'}} \frac{\gamma_S}{\omega_S}$$

in situ observation

Kinetic simulations

Before STEREO

- **spectrum**: simultaneous observations of high (\sim fp) and low (\sim 1kHz) frequency waves
- **waveform**: only access to the high frequency (too short time waveforms) - only two axes

phase information lost

phase information

Before STEREO

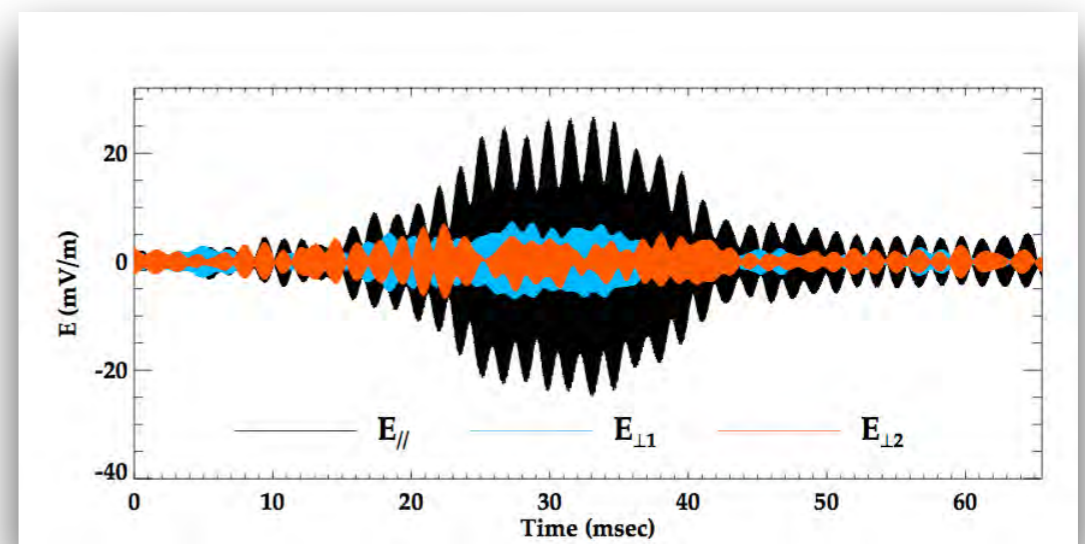
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phase information

From STEREO waveforms we can access

- **simultaneously** to the high & low frequency signals
- 3D structures of the waves (3 antennas)



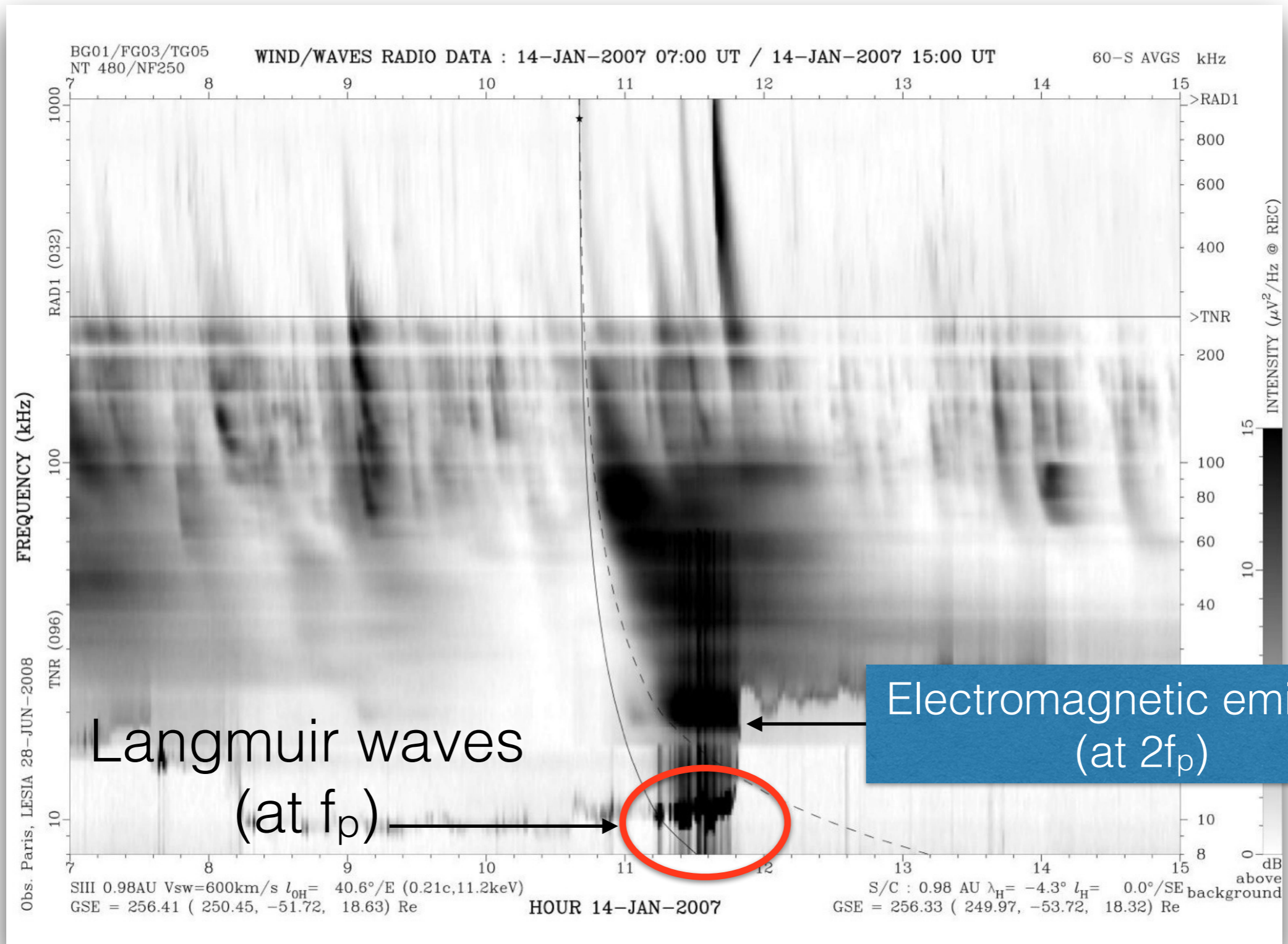
A case study

Langmuir decay & wave coupling process
confirmed from
observations & simulations

Henri P., Briand C., Mangeney A., JGR, A03103, 2009

Henri P., Califano F., Briand C., Mangeney A., JGR, A06106, 2010

STEREO Jan. 2007



Check of the fundamental laws:

1 - Energy & Momentum conservation

Theory:

$$\omega_L = \omega_{L'} + \omega_{IAW}$$
$$\vec{k}_L = \vec{k}_{L'} + \vec{k}_{AW}$$

Observed:

$$f_L^{Doppler} = f_L + \frac{\vec{k}_L}{2\pi} \cdot \vec{V}_{sw}$$

The frequency condition for resonance

$$f_L = f_{L'} + f_{IAW}$$

checks simultaneously the two conservation laws
(energy & momentum)

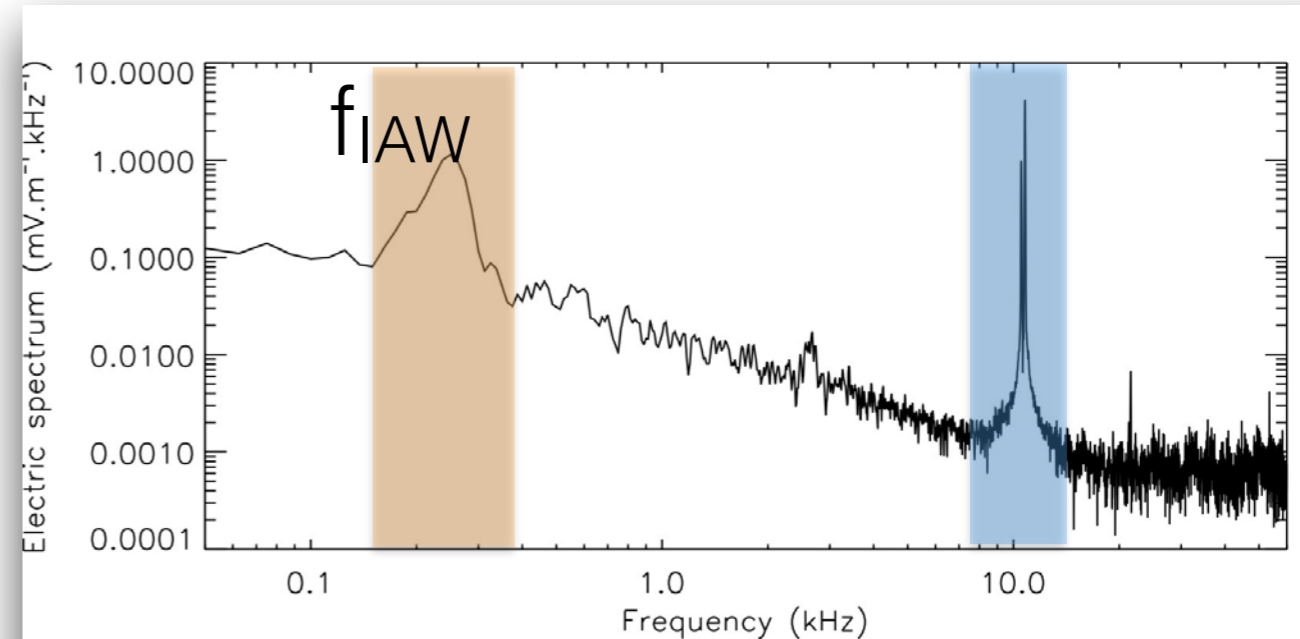
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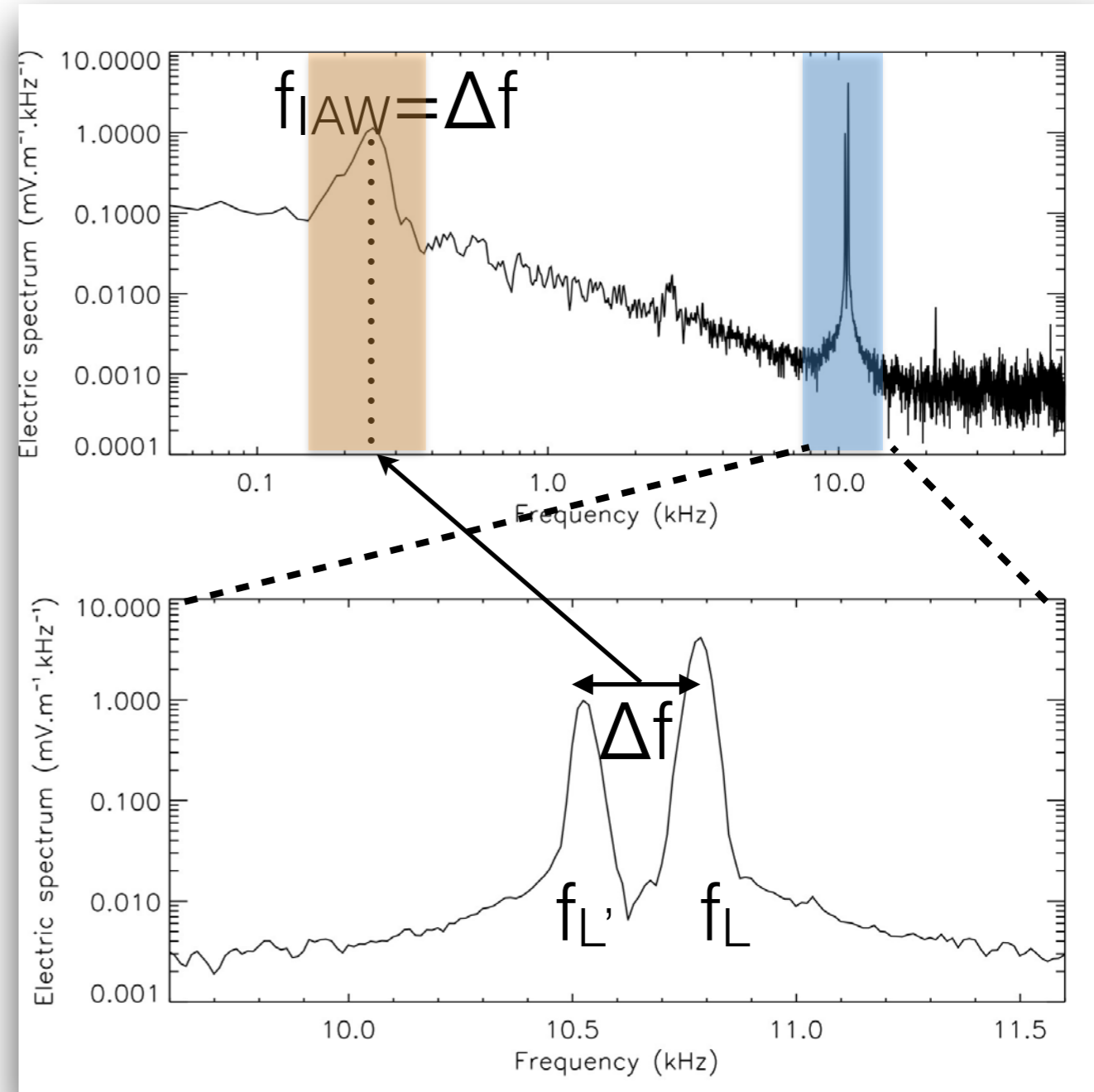
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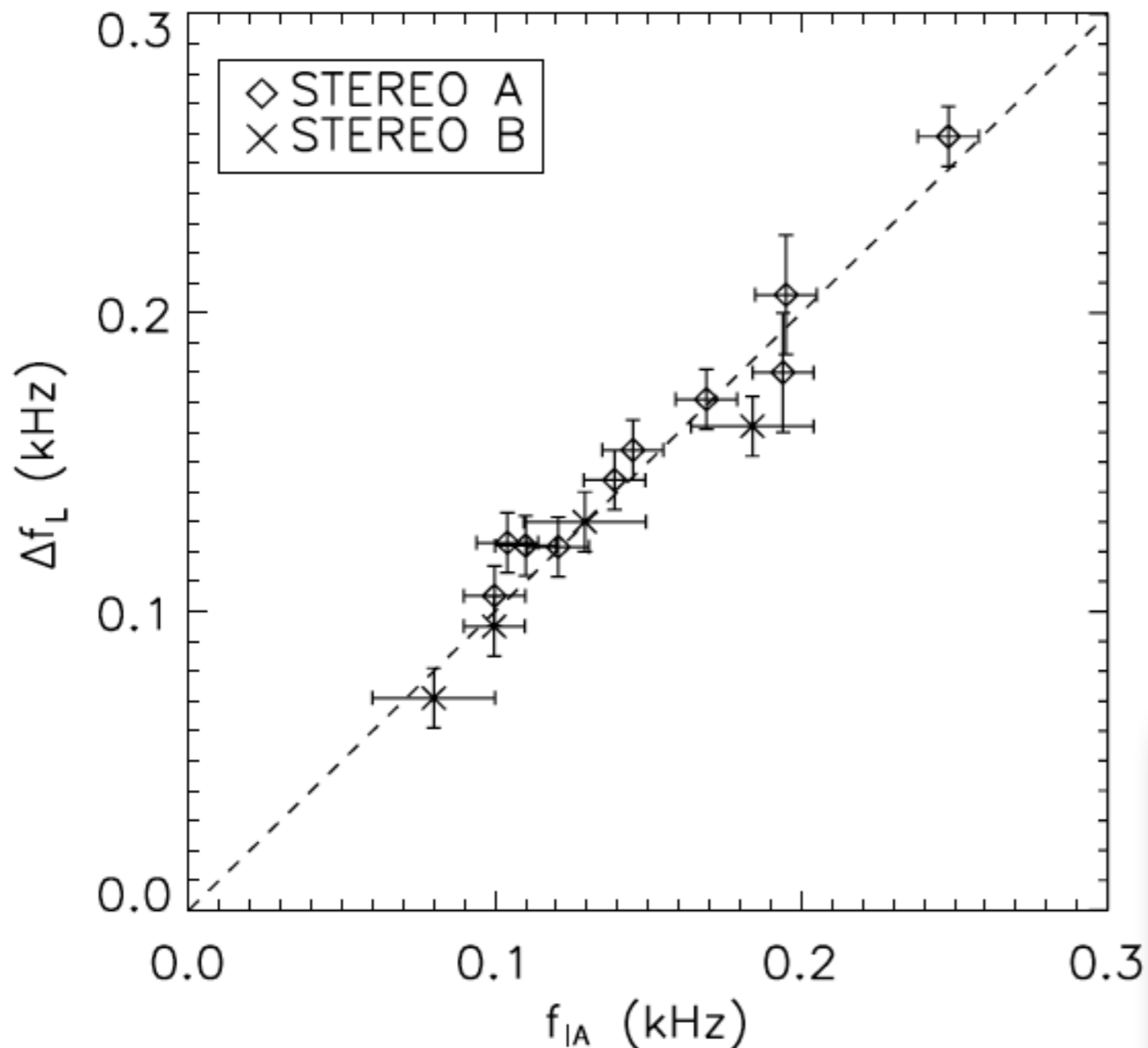


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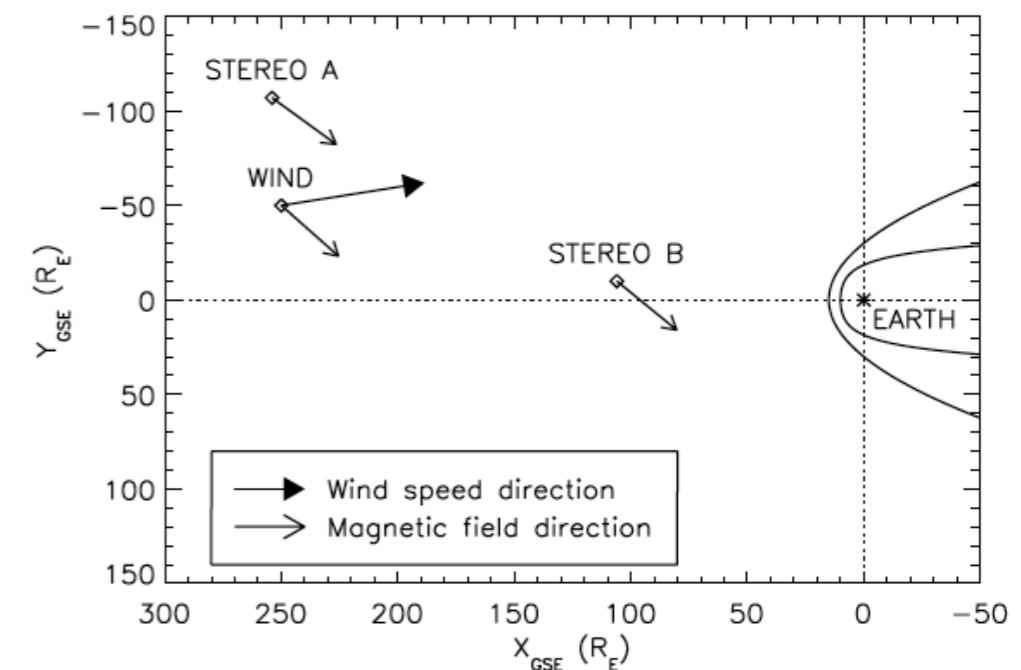
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Check of the fundamental laws: 1 - Energy & Momentum conservation



From all the Langmuir events of the Type III burst of January 2007 from STA & STB (still close)



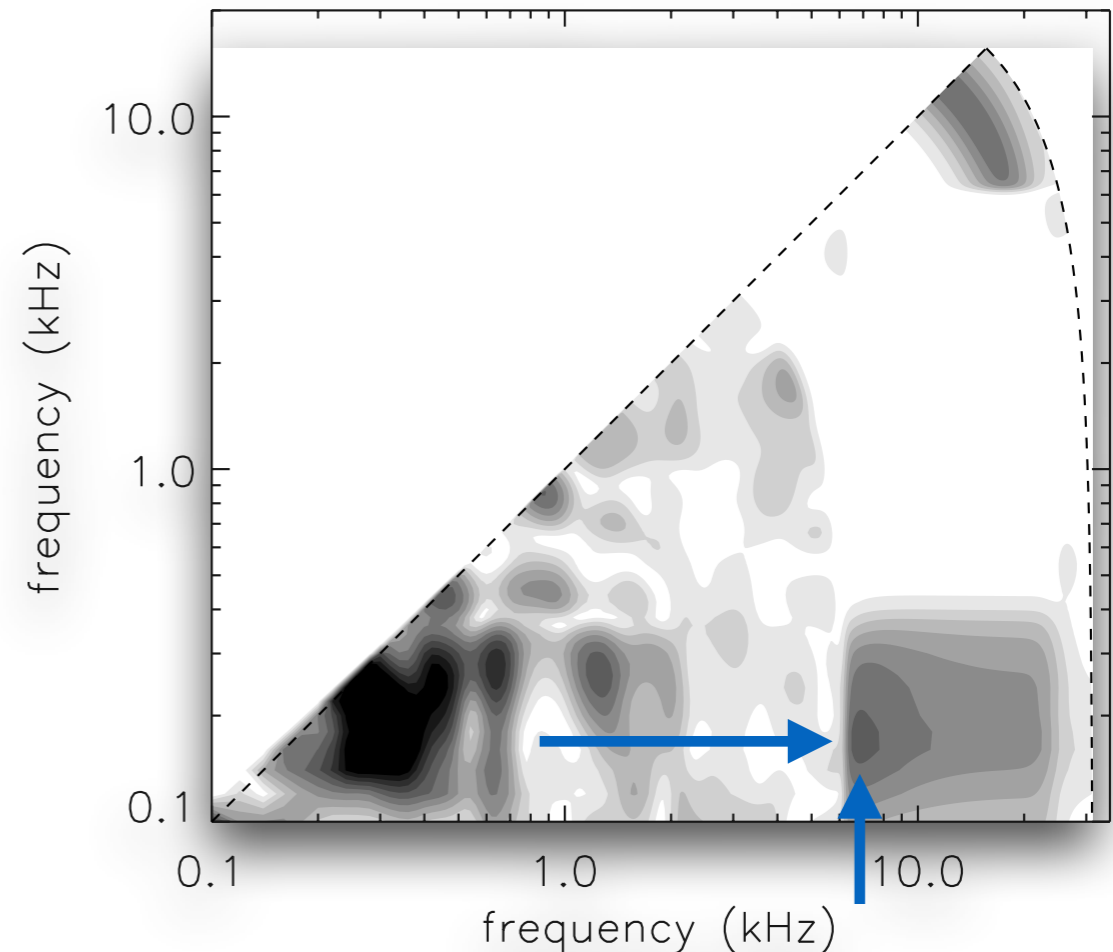
Fastest electrons: $4 \cdot 10^5 \text{ km/s} \sim 0.13c$

Check of the fundamental laws: 2 - Phase locking

Bicoherence analysis

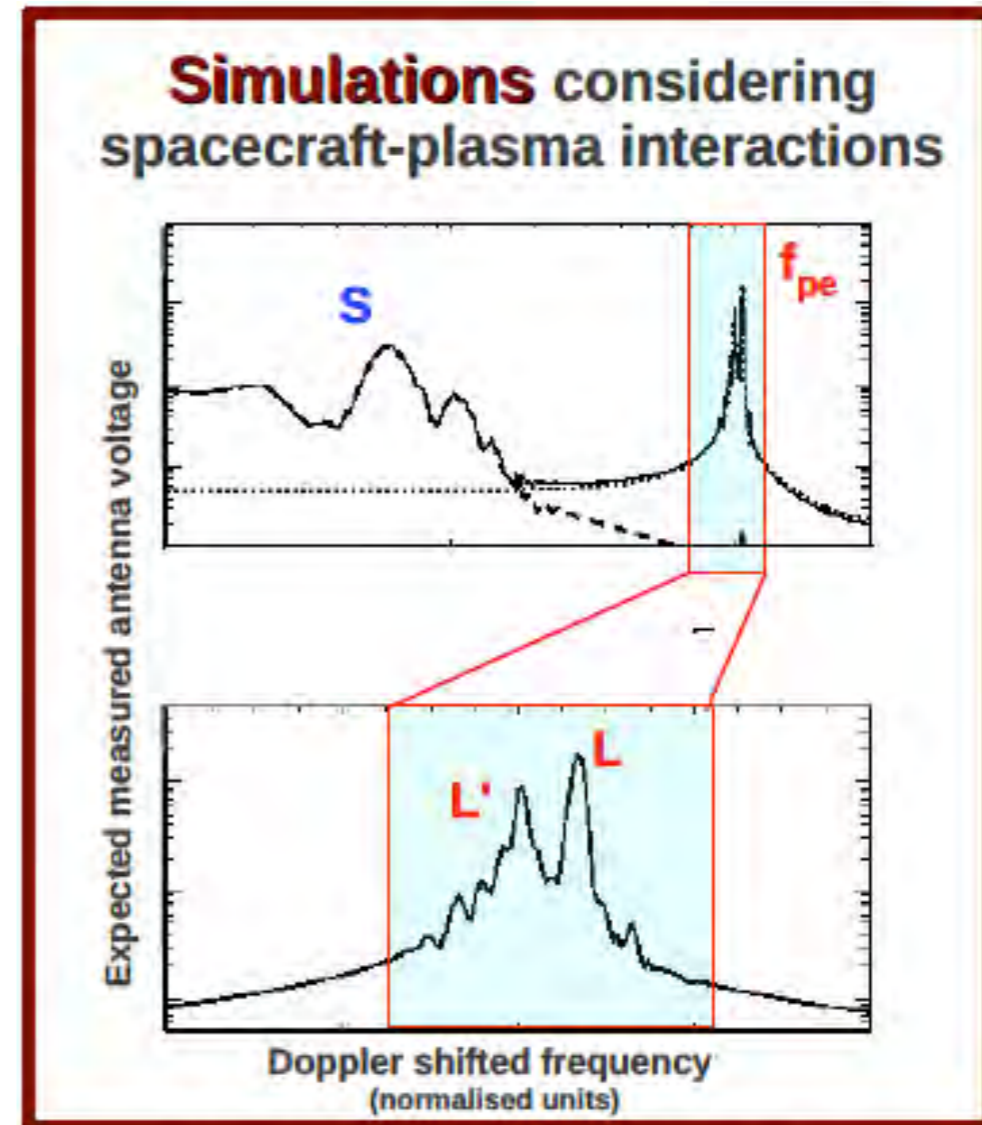
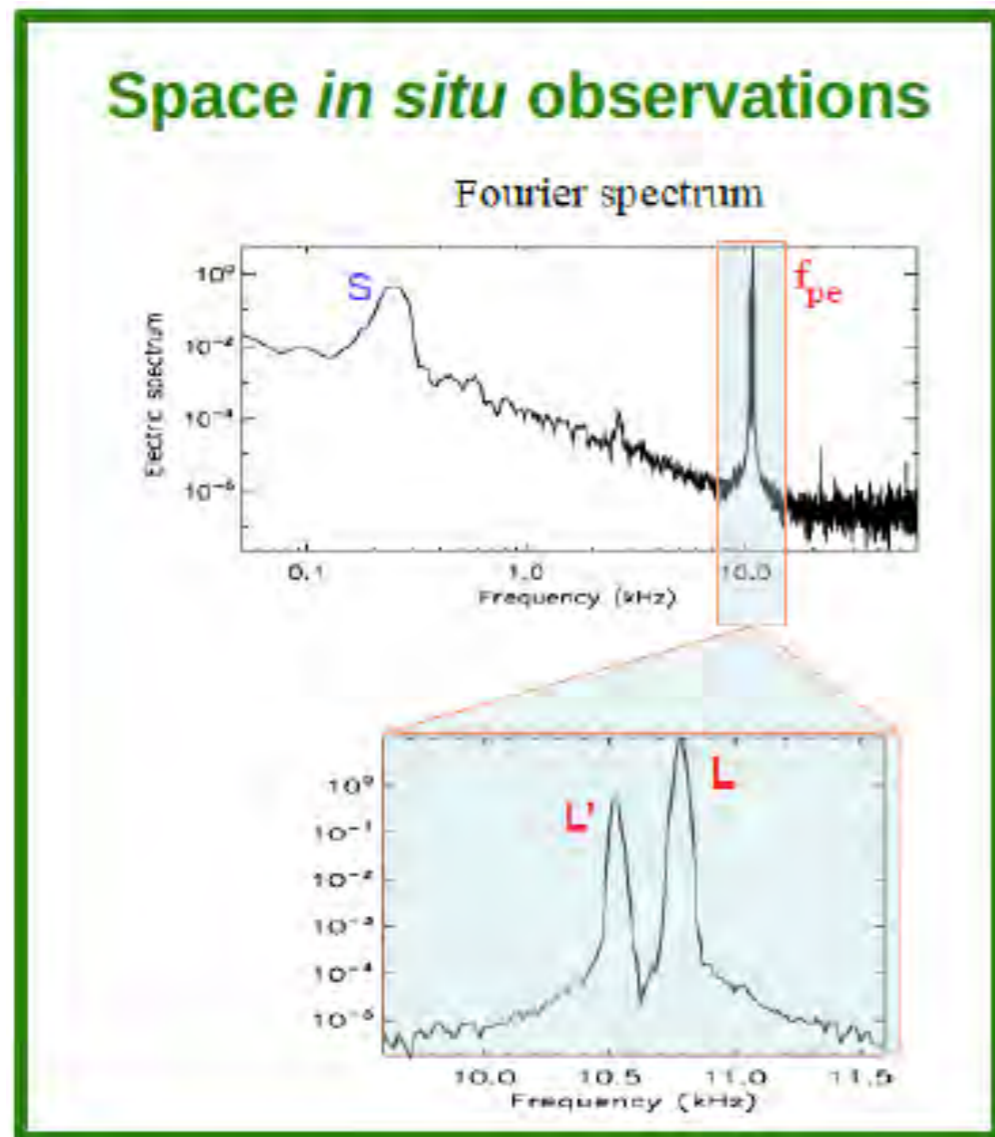
- measures the stationarity of the relative phase of the waves
- **statistical method**: must be applied to a large set of waveforms

From **14** spectra of a same Type III
simultaneously analyzed



Amplitude threshold reached ? What the simulation says ...

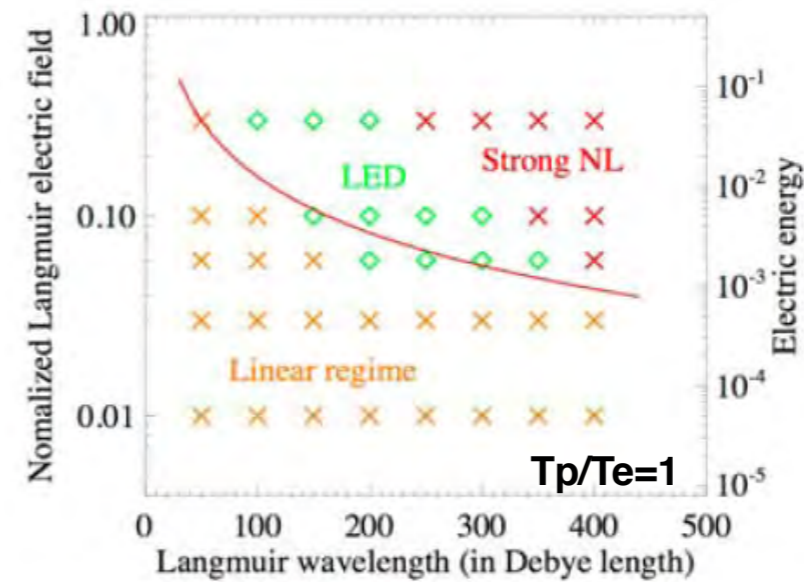
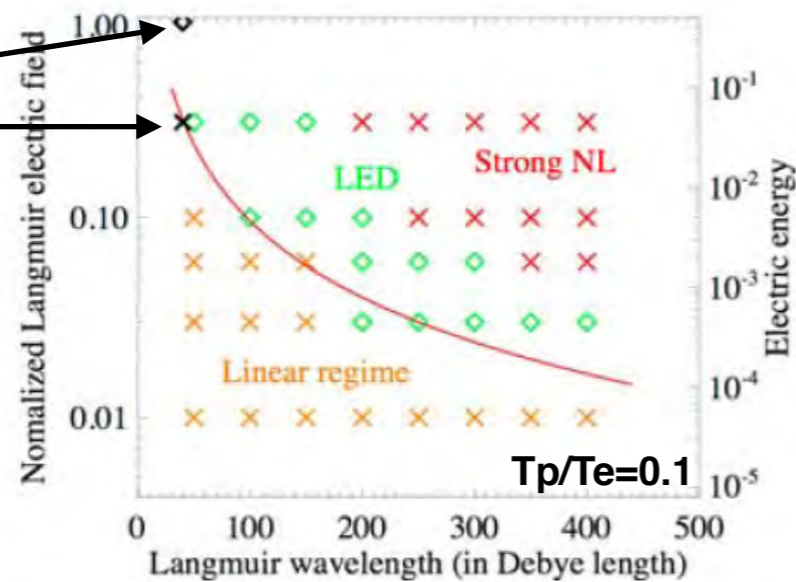
Vlasov-Poisson
1D-1V electrostatic, Periodic conditions



[Henri et al., Solar WIND 12, 2010]

Langmuir Electrostatic Decay: new threshold for non monochromatic waves (wave-packets of width Δ)

Umeda (2007)
Umeda & Ito (2008)



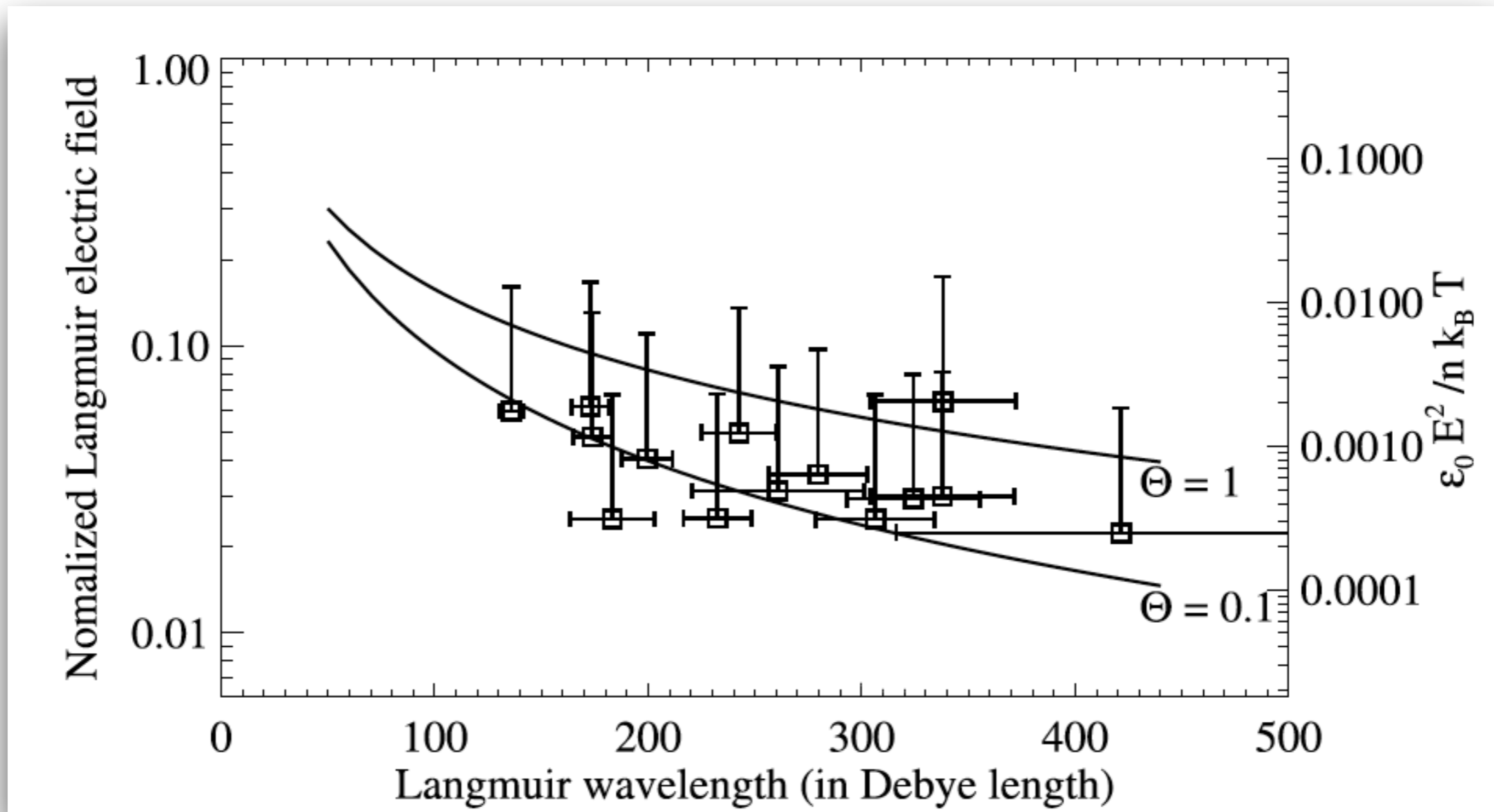
$$E_{LED}^{thres} = \left(\frac{6k_L^{1-\beta}}{\Gamma\Delta} \right)^{1/\alpha}$$

Growth rate

$$\gamma_{LED} = \Gamma E_L^\alpha k_L^\beta$$

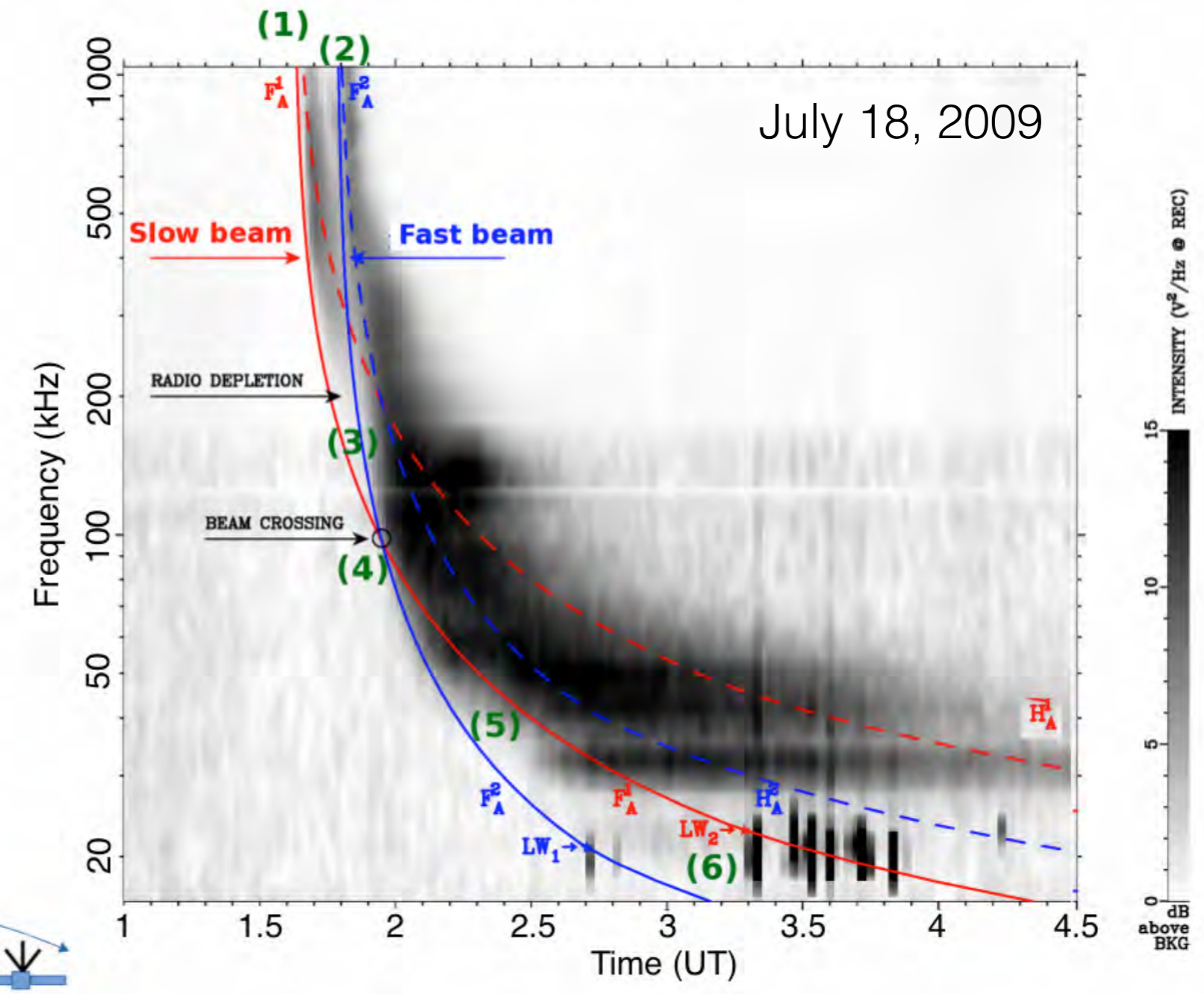
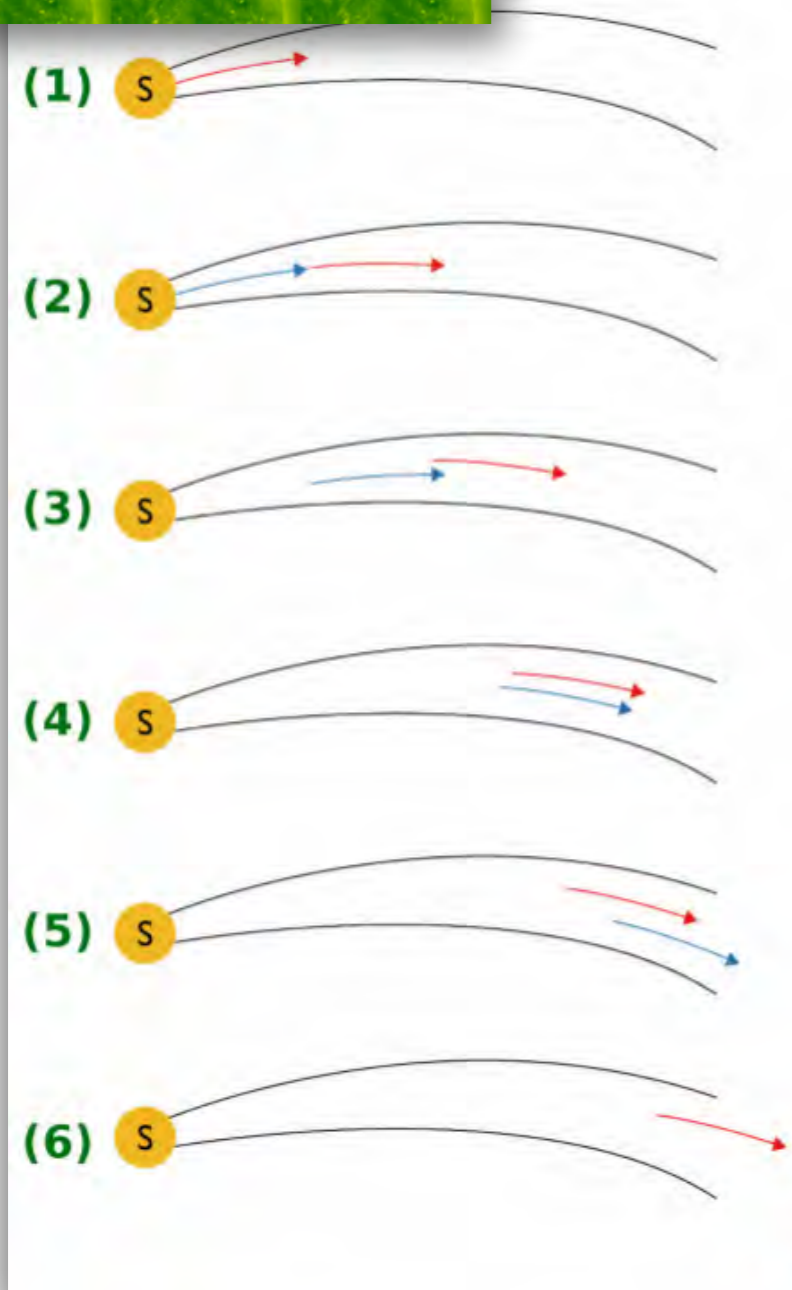
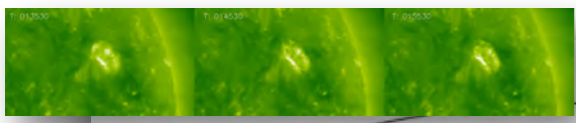
	Γ	α	β
$T_p/T_e = 0.1$	0.026	1.11	0.59
$T_p/T_e = 1$	0.025	1.82	0.30

Langmuir Electrostatic Decay: Observations fit between the thresholds



A second case

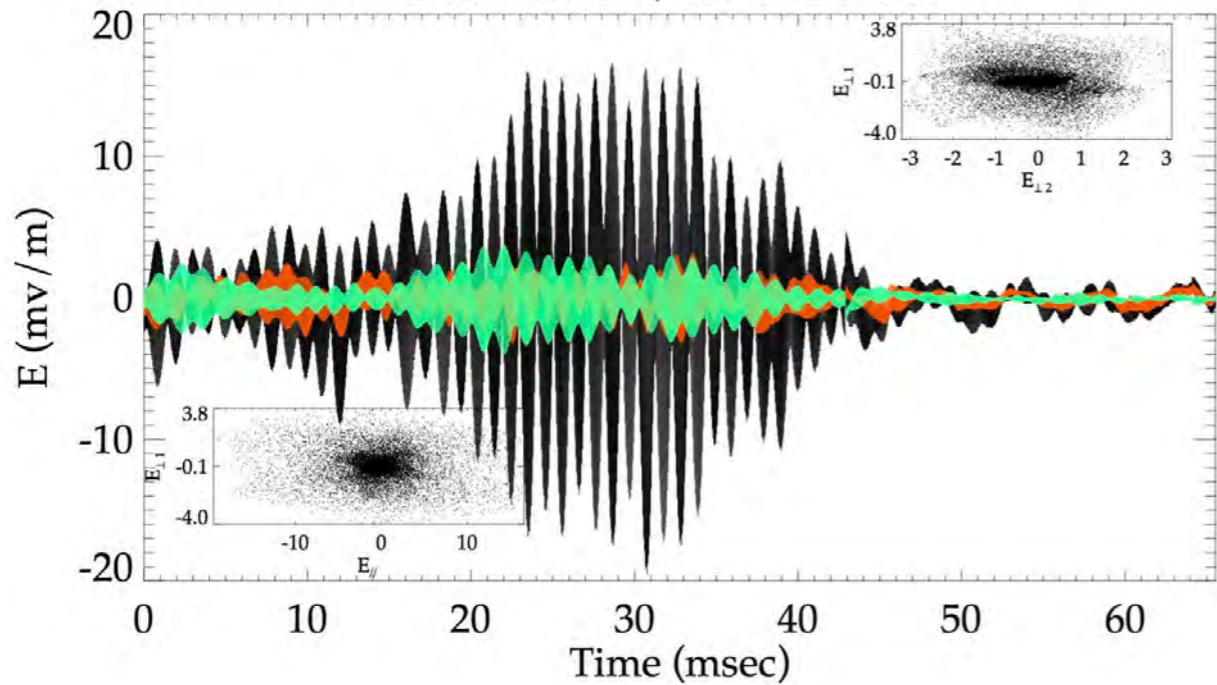
Parametric (threshold) process
Phase coherency importance



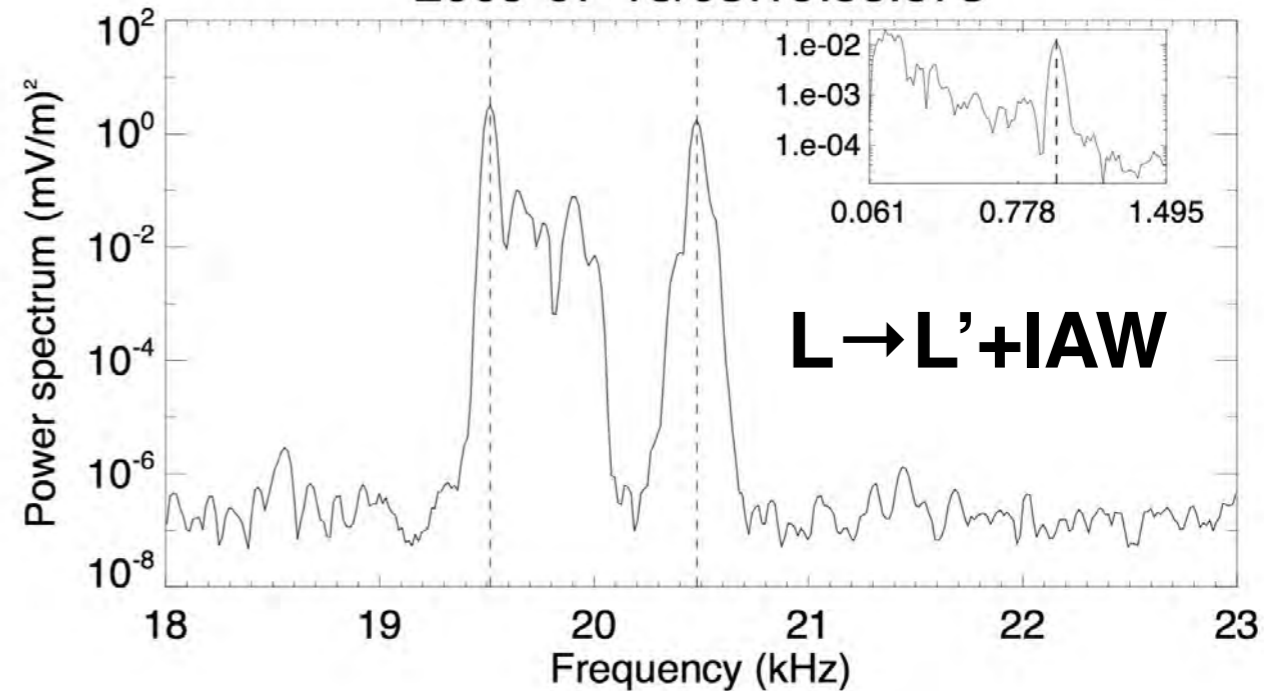
- 1° Start of a first beam (01:38) - LW detected at 03:17; slow beam (0.09c)
- 2° Start of a second beam (01:47) - LW detected at 02:42; fast beam (0.16c)

At the time the two beams intersect (4), decrease of the Type III intensity

2009-07-18/03:19:55.573

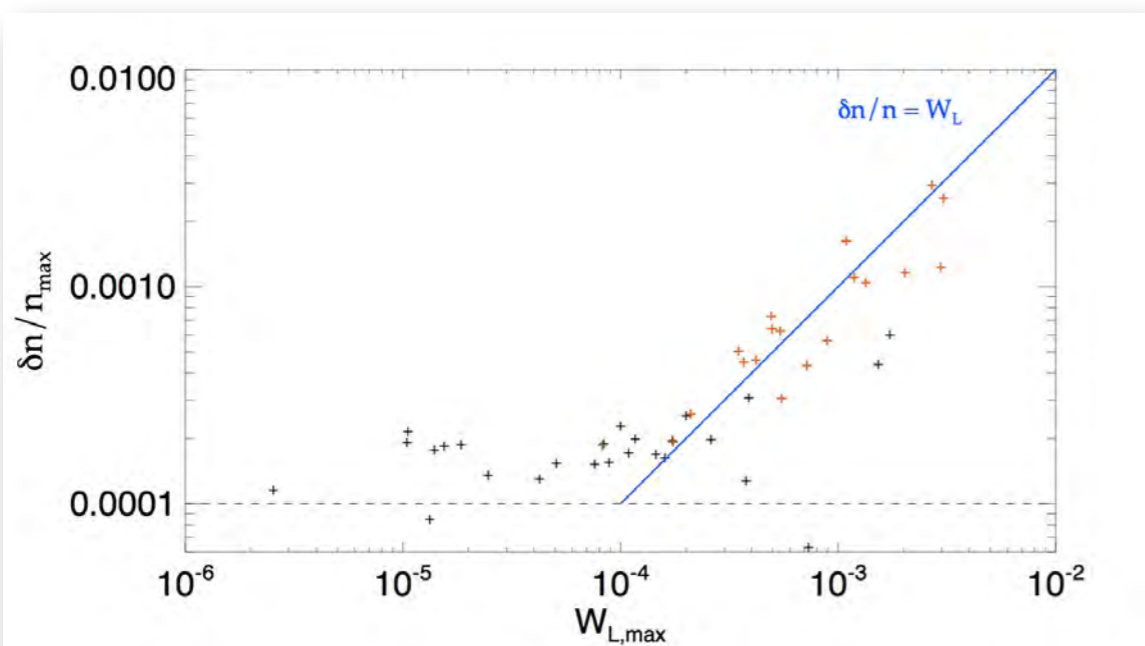


2009-07-18/03:19:55.573



$$V_b \approx \frac{2f_p V_{SW} |\cos(\theta)|}{f_{IA}}$$

Compatible with the EM observations



$$\delta n/n \sim \epsilon_L / \epsilon_{th,e}$$

Level of density fluctuation as expected when associated with Langmuir decay

In situ signatures compatible with Langmuir decay

Vlasov Simulations

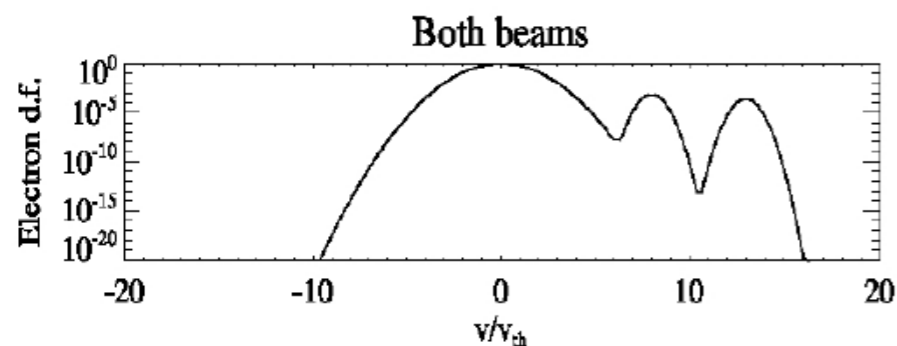
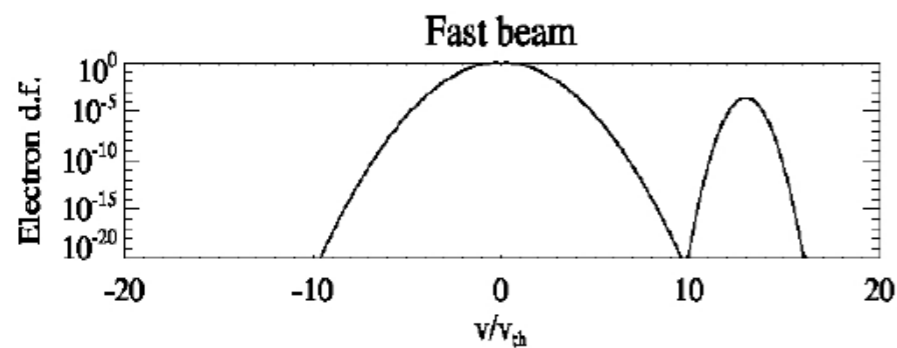
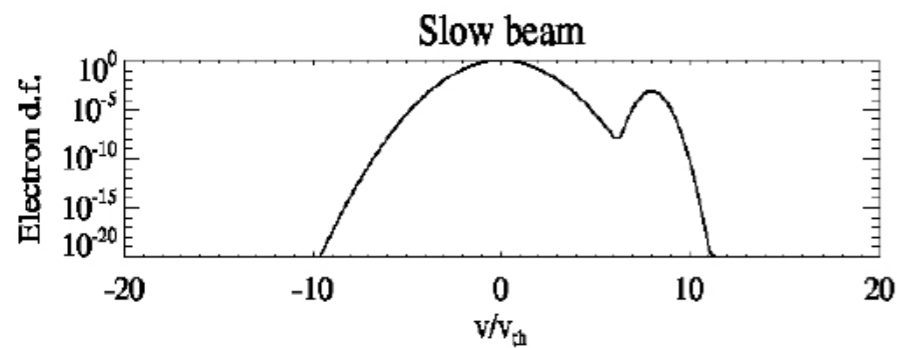
Box size: $L \sim 10,000$ Debye lengths; Periodic boundary conditions;

Initial conditions: - Maxwellian distribution functions,

+ electron beams :

Beam speed	Beam thermal speed	Beam density
$8 V_{th}$	$0.5 V_{th}$	$1.5 \cdot 10^{-2}$
$13 V_{th}$	$0.5 V_{th}$	$0.5 \cdot 10^{-2}$

**Suppose initially
*homogeneous plasma***

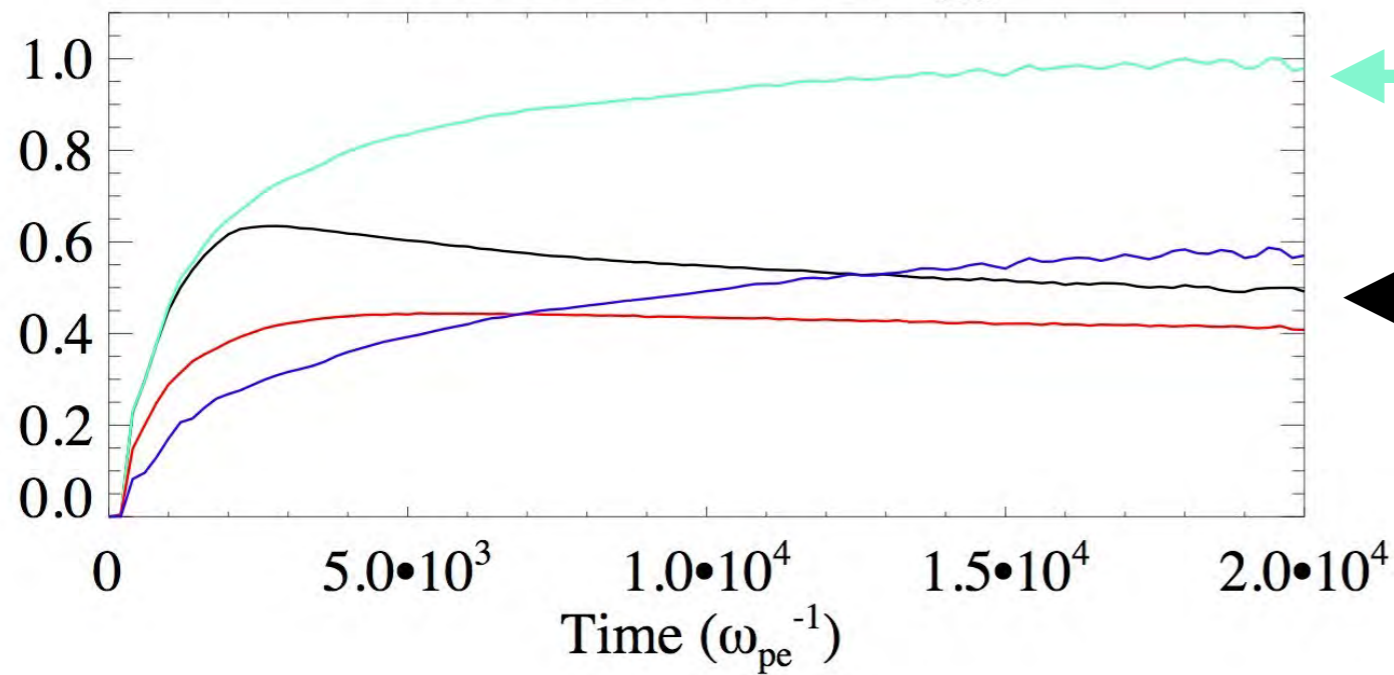


Slow beam only

Fast beam only

Two beams

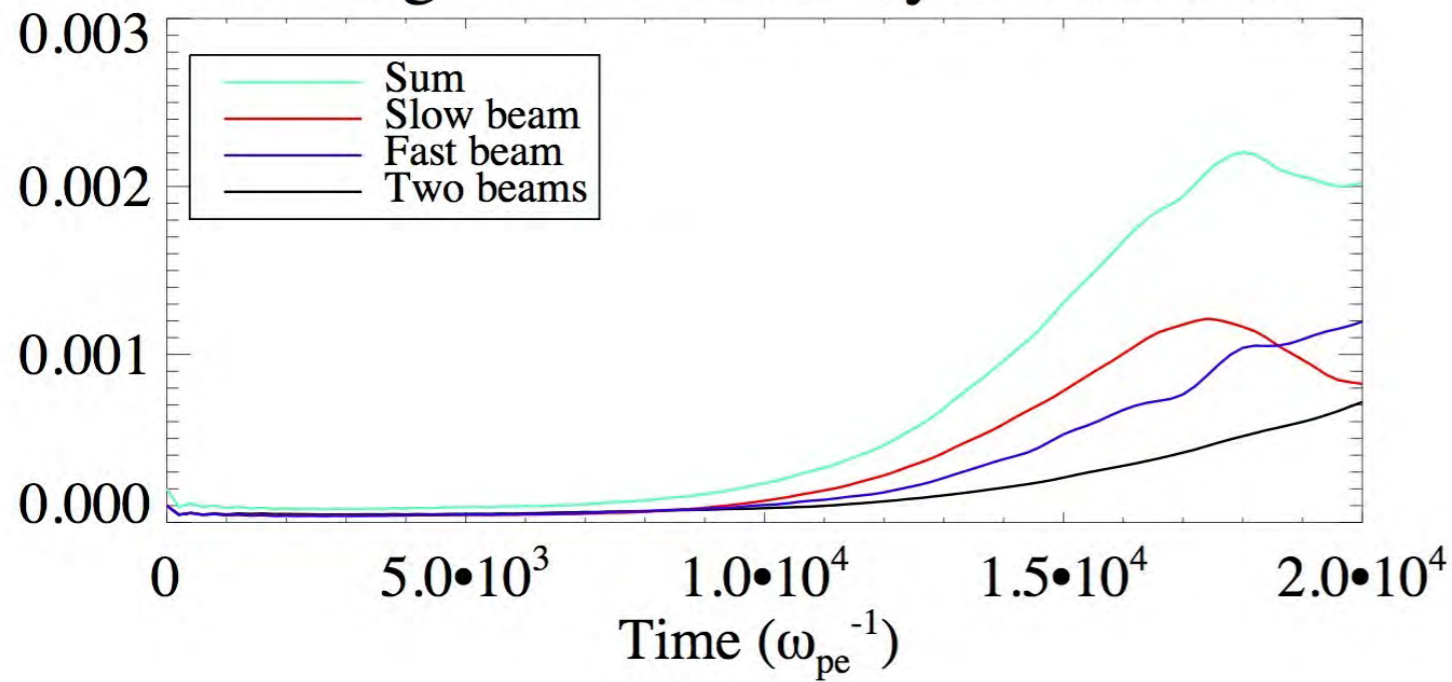
Total electric energy



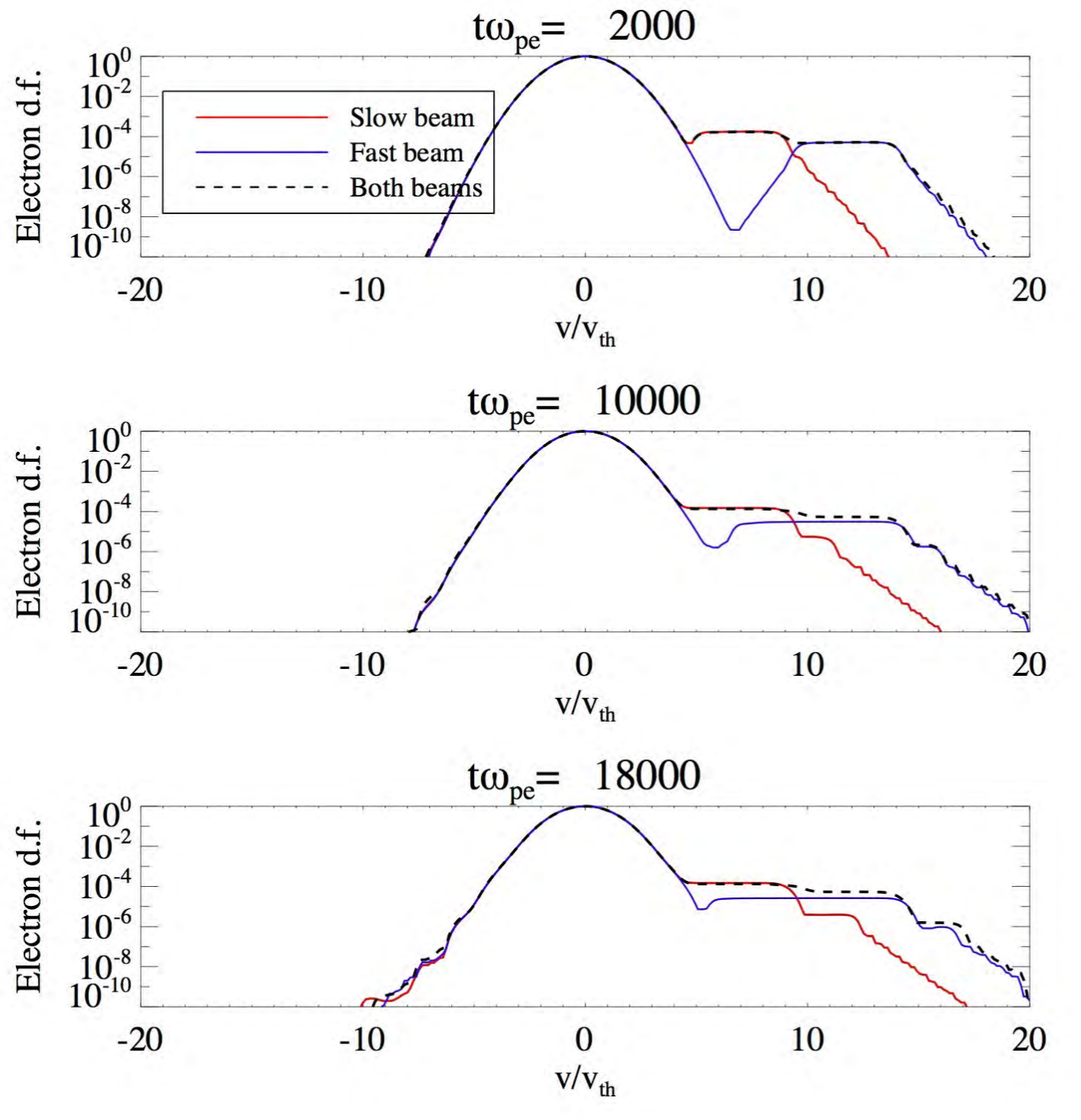
Sum of two isolated beams

Interaction of two beams

Average level of density fluctuations



*decay-produced density fluctuations:
proxy of the efficiency of the Langmuir decay*



Screening of the fast beam by the presence of the slow beam



Less electric energy conversion



Lower level of LW

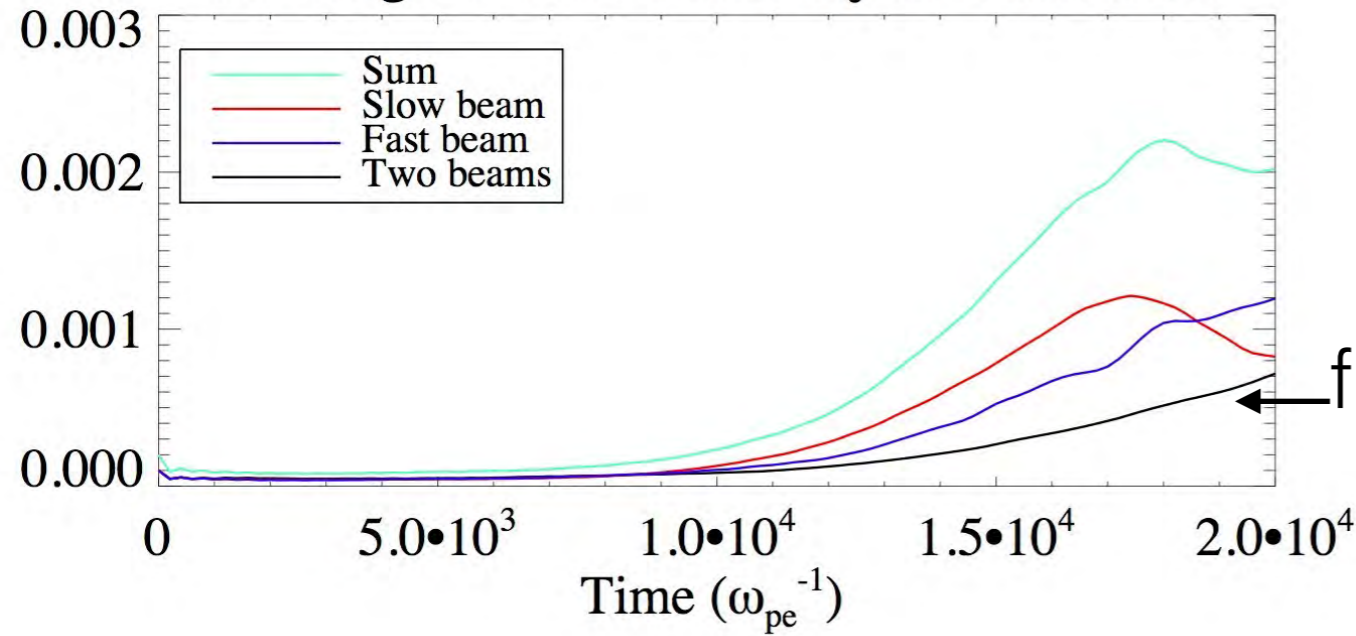


Lower level of EM

Beam-driven Langmuir waves generation is less efficient:

whatever the ES to EM conversion mechanism, as long as it depends on the beam-driven Langmuir waves amplitude, production of the EM drops

Average level of density fluctuations



With two beams, the level of density fluctuations is *always* smaller than the contribution with individual beams

⇒ Not only the fast-beam-driven LW are affected by the presence of the slow beam.

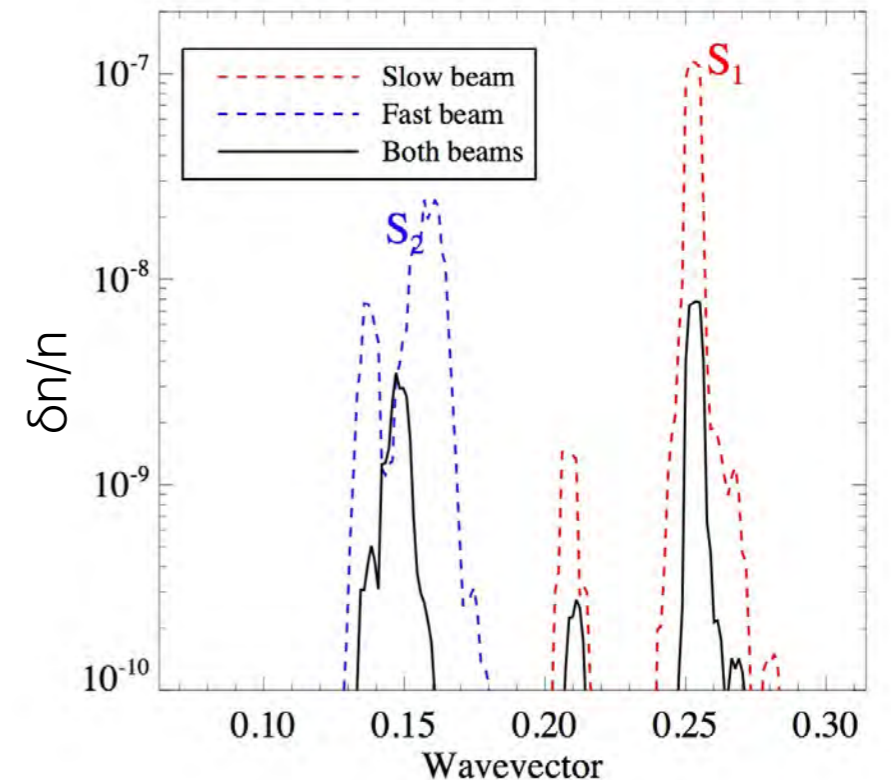
slow-beam-driven LW generation is also disturbed

1° Threshold process

even a small decrease of the mother LW energy can make the process inefficient

2° Coherent process

Phase beating between 3 waves. Ease if $k_{IAW} \neq k_{L,L'}$.
With 2 beams, larger spectrum of k : reduction of the phase locking - seen in the $\delta n/n$ proxy



A composite image featuring a large, detailed view of the Earth's blue and white surface on the left side. In the center, a bright sun is partially obscured by the Earth's horizon, creating a lens flare effect. The background is a dark, starry space filled with numerous small, distant stars.

Conclusions

Strongest solar radio emissions
depend on electrostatic waves
produced by electron beams

First time the fundamental laws
can be checked.

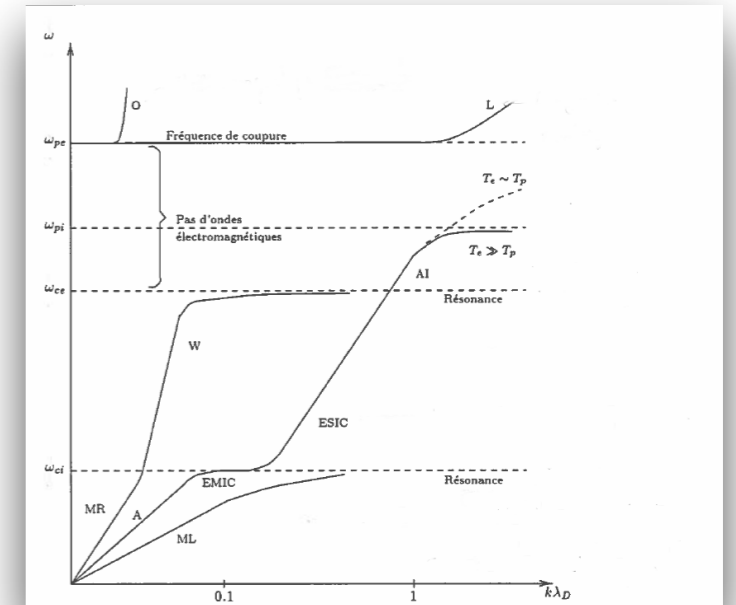
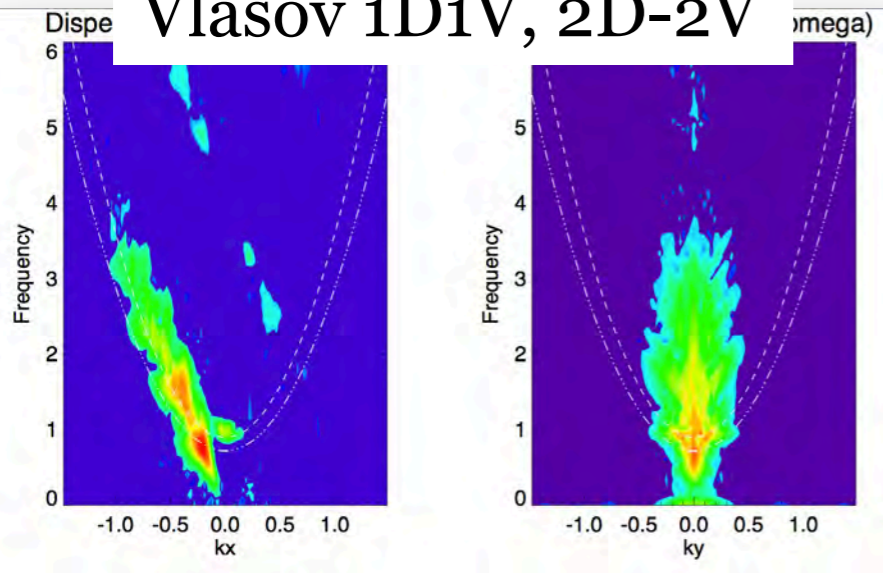
Not only Langmuir wave amplitude
is of importance but also the
coherence between the waves

Langmuir decay threshold
compatible with observations
(new theoretical threshold)

A good way to transfer energy
from high to low frequencies

NEXT ...

Vlasov 1D1V, 2D-2V



Parametric decay in 2D, inhomogeneous plasma

Coupling with lower frequency waves (transfer of energy - turbulent cascade)

Labs experiments (laser facilities)

Laser plasma interaction (Stimulated Raman Scattering)

$$T_{fp} \rightarrow L + LF$$

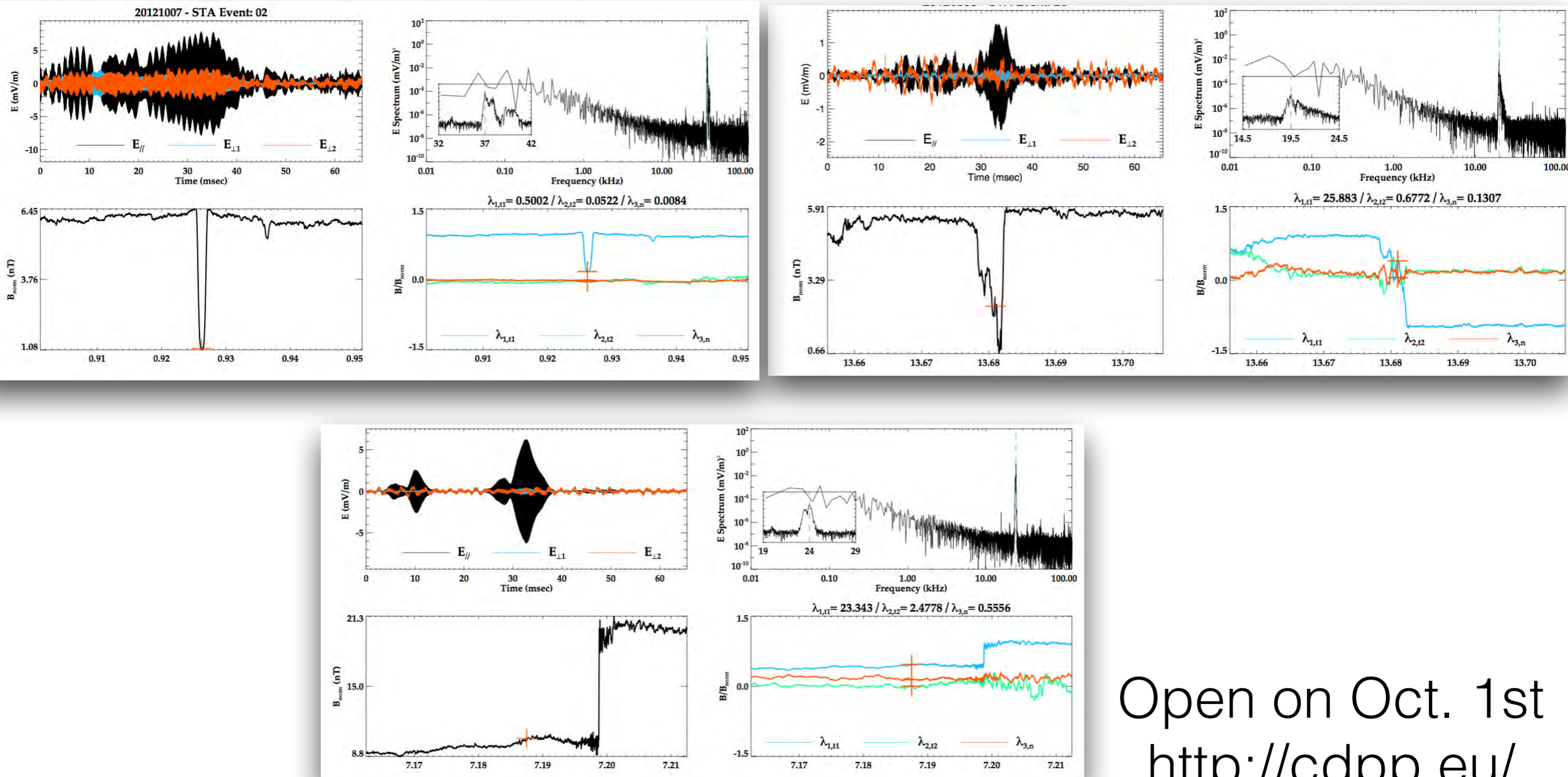


	$\rho_e \text{ (cm}^{-3}\text{)} / f_p$	T_p/T_e	$T_e \text{ (keV)}$	$k\lambda_d$
ICF(1micron)	$10^{12-19} / 900\text{MHz}$	0.1-1	0.1-1	0.01-0.5
Space (1AU)	1-10 / 10 - 30kHz	0.1 - 1	10^2	< 0.05

T_p/T_e and $k\lambda_D$: comparable

CDPP to access a database with ALL Langmuir observations of STEREO (2007-2014)

CDPP : Centre de Données de Physique des Plasmas / Center of Plasma Physics Data



Open on Oct. 1st
<http://cdpp.eu/>

A dramatic, low-angle shot from a surfer's board camera, looking through the barrel of a massive, curling wave. The water is dark and turbulent, with a bright, golden sunset visible through the opening of the wave. The sun is low on the horizon, casting a warm glow across the scene. The text "PAU & MAHALO" is overlaid in the center of the image.

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<http://www.dailymail.co.uk/news/article-2578228/Surfer-ditches-board-camera-capture-precise-moment-biggest-waves-break-positions-right-them.html>