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Parametric Tolerances for the WIYN Wide Field Corrector

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Introduction

The optical design of the WIYN telescope calls for corrective optics to be placed in the light path to one of the Nasmyth foci. The Wide Field Corrector (WFC), which was designed by Charles Harmer, is a two element lens 50 cm in diameter. Both elements are of fused silica and have all-spherical optical surfaces. When installed in the telescope the WFC will provide image correction for a one degree diameter field of view with a bandpass from 330 nanometers to 1.53 microns.

The location of the WFC is shown relative to the telescope reflective optics in figure 1. Figures 2 and 3 show the WFC lens elements, and figure 4 shows spot diagrams for rays traced through the corrector. These spot diagrams show the combined images of five wavelengths of light (at 330, 365, 436, 852 and 1529 nanometers) at five locations in the field (on axis and at 0.125, 0.25, 0.375, and 0.5° off axis). The RMS and geometric (ie 100% encircled energy) image radii are listed in the figure.

Parametric tolerances for the WFC consist of the radii of curvature of the four surfaces, lens thicknesses and wedges angles. Full fabrication tolerances must also consider the quality of the optical surfaces as departures from ideal surfaces.

Parametric Tolerancing Procedure

The Zemax ray tracing program was used to determine the effect of changes in each of the eight fabrication parameters. To do this the parameter was changed a small amount, then the Zemax optimization package was used to determine a new best position along the telescope optical axis for the WFC. The radius of curvature of the focal plane was also optimized.

During the optimization the total track length from the primary to the secondary and from the secondary to the focal plane was held constant. In all cases a corrector shift on the order of 5 mm or less was sufficient to largely compensate for perturbations to the lens parameters.

In each case the nominal design was used as a starting point for the optimization. This was done for repeatability so that any particular optimized configuration would not be influenced by the optimization history.

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Perturbations in six of the parameters (the four radii and two thicknesses) maintain an axisymmetric optic, while the two wedge angle result in non-symmetric optics. For the symmetrical
cases, after optimization the RMS image radii were recorded for five filed positions in five colors
as described earlier. The results are presented graphically in figures 5 and 6, which show the rms
image radius for the five field positions and the average image radius as a function of variation
in surface radius of curvature. Note that in all cases the image diameter at the edge of the field
(circular tic marks) is most influenced by parametric variations, while the average value is
relatively unchanged for small departures from the design values.

For the case of the two wedge angles, spot diameters were recorded at 17 points in the one degree field arranged on orthogonal axes.

The influence of each parameter on image size and on focal plane curvature was determined from the data. In cases where the slope was not symmetric with respect to positive or negative departures from the nominal design, the larger of the two influences was recorded. These were used to assign tolerances to each parameter.

The tolerances were balanced to assign the largest allowance to parameters which had the strongest influence. Other parameters such as wedge and thicknesses, which are relatively easy to hold, were set to values that can be met with standard shop practices.

Parametric Tolerances

The influence functions and tolerances are given in table 1 attached. As a "spot" check, the corrector design was perturbed by randomly setting all the lens parameters to positive or negative extremes of the tolerance range. Two such random settings yielded average image size of 9.65 and 9.244 microns RMS radius (0.153 and 0.146 arcseconds FWHM respectively). This compares to a nominal design value of 8.984 microns RMS radius (0.142 FWHM). This constitutes an image degradations of 0.007 and 0.004 arcseconds FWHM, well within the image error budget allowance of 0.04 arcseconds FWHM.

Table 1: Summary of influence functions & tolerances for WIYN Wide Field Corrector May 11, 1993

Total image allowance: 0.04 fwhm = 0.0472 RMS = 2.5252 Microns RMS Radius

Focal surface allowance: +/-.1 mm change in sagitta = +/- 155 mm change in radius

8 parameters for fabricating the two lenses:

	Wedge 3-4	Wedge 1-2	772	T1	R4	ፎ	R2	R:			Parameter
	0	0	20	20	519.991	599.77	1954.931	1424.044		Value	Nominal
	deg	deg	mm	mm	mm	mm	mm	mm			Unit
19.454	9.77	0.55	1.196	0.016	4.2	3.3	0.112	0.31	(microns)	influence	Image
100.00%	50.22%	2.83%	6.15%	0.08%	21.59%	16.96%	0.58%	1.59%		influence	% of total
	0.01	0.01	0.2	0.2	0.4	0.5	w	2			Tolerance
	0.09770	0.00550	0.23920	0.00320	1.68000	1.65000	0.33600	0.62000	(microns)	Contribution	Image
	0.0015	0.0001	0.0038	0.0001	0.0266	0.0261	0.0053	0.0098	FWHM	Contribution	Image
	0	0	62	-30	124	95	14	-18	Influence	Radius	Focal
	0	0	12.4	ტ	49.6	47.5	42	-36	Contribution	Radius	Focal

RSS (microns RMS radius) 2.47164

Arcseconds RMS 0.04620 FWHM 0.03915

of focal surface radius (mm) 89.2534

RSS

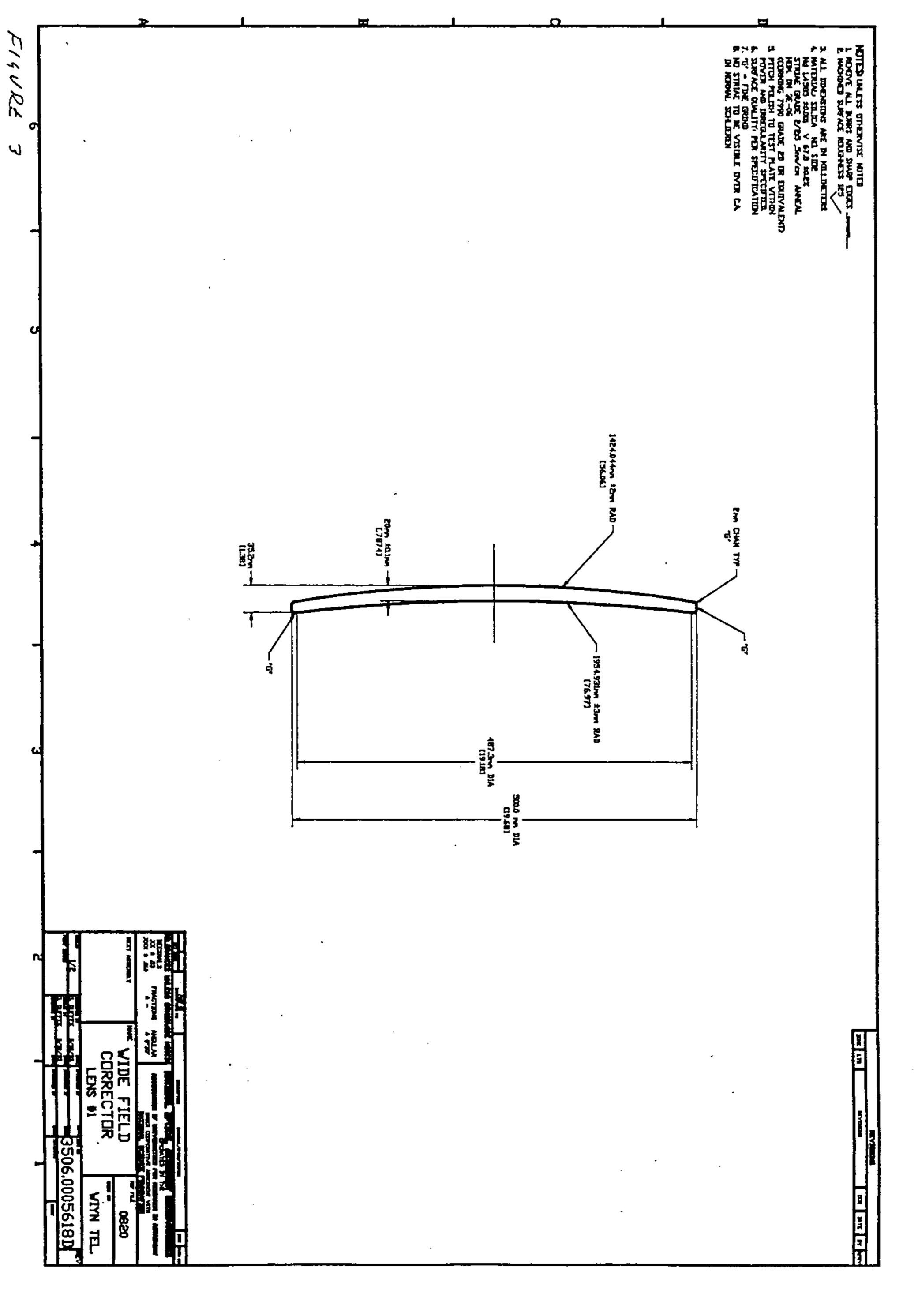
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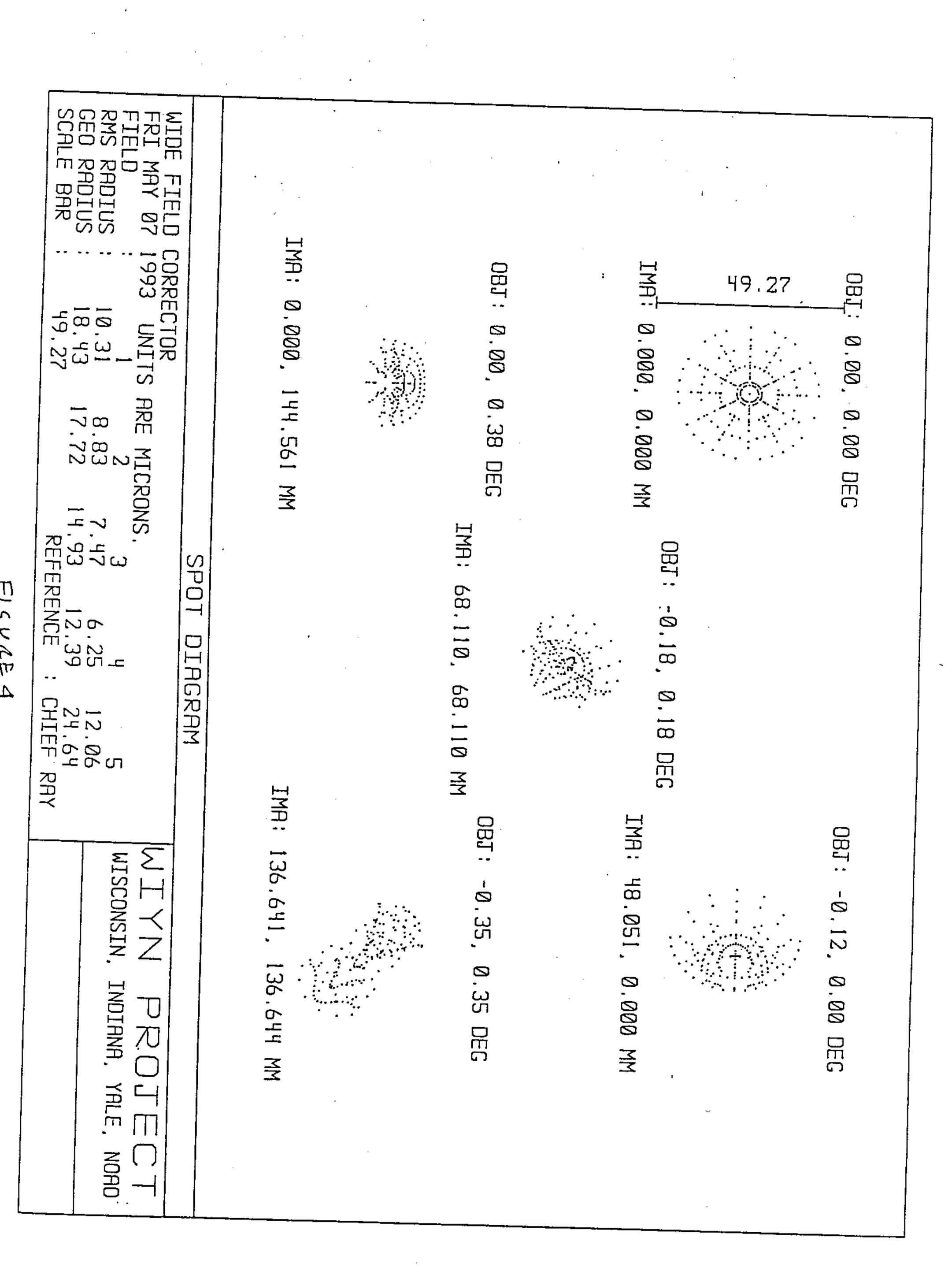
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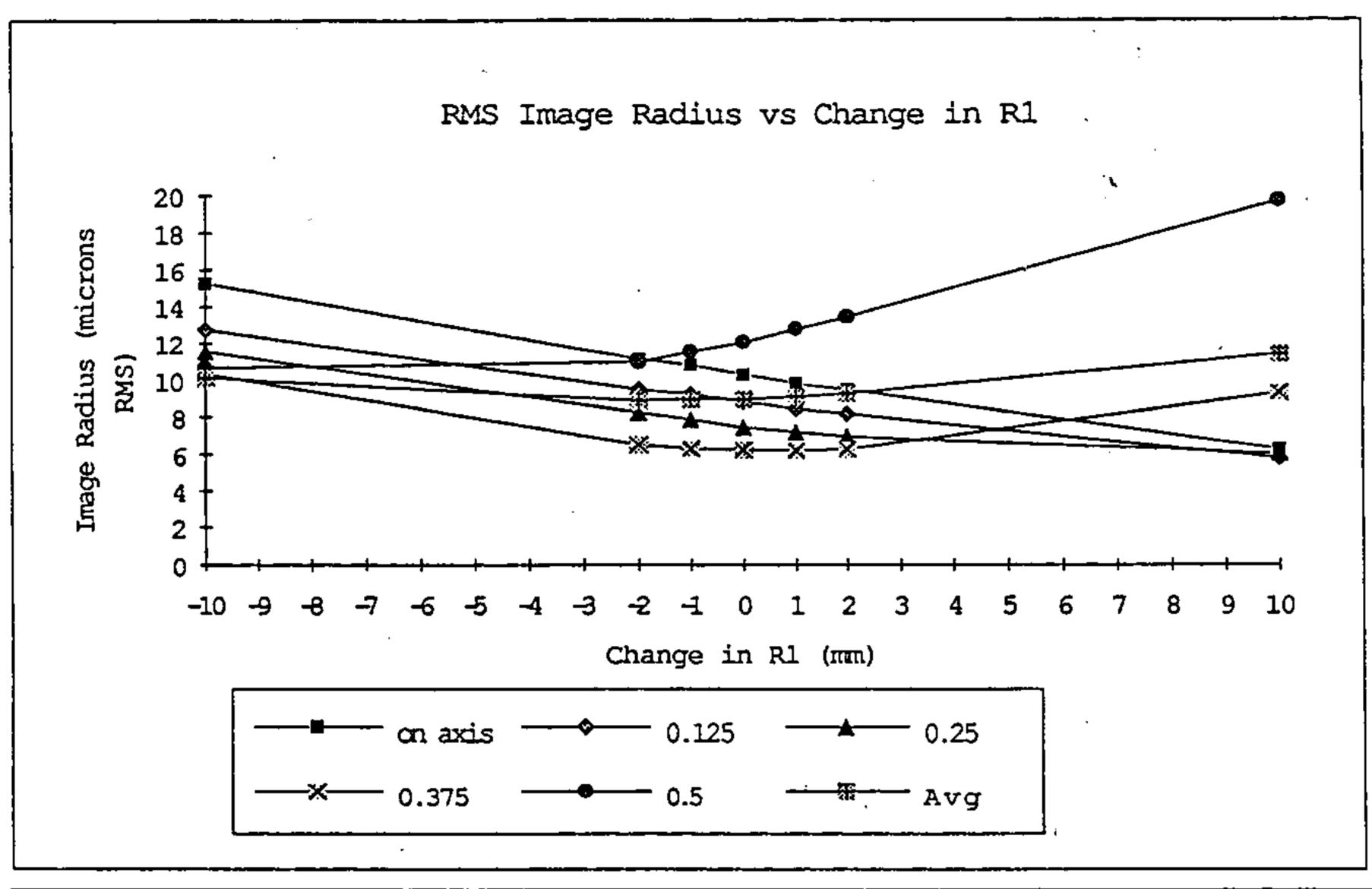
N 3.4

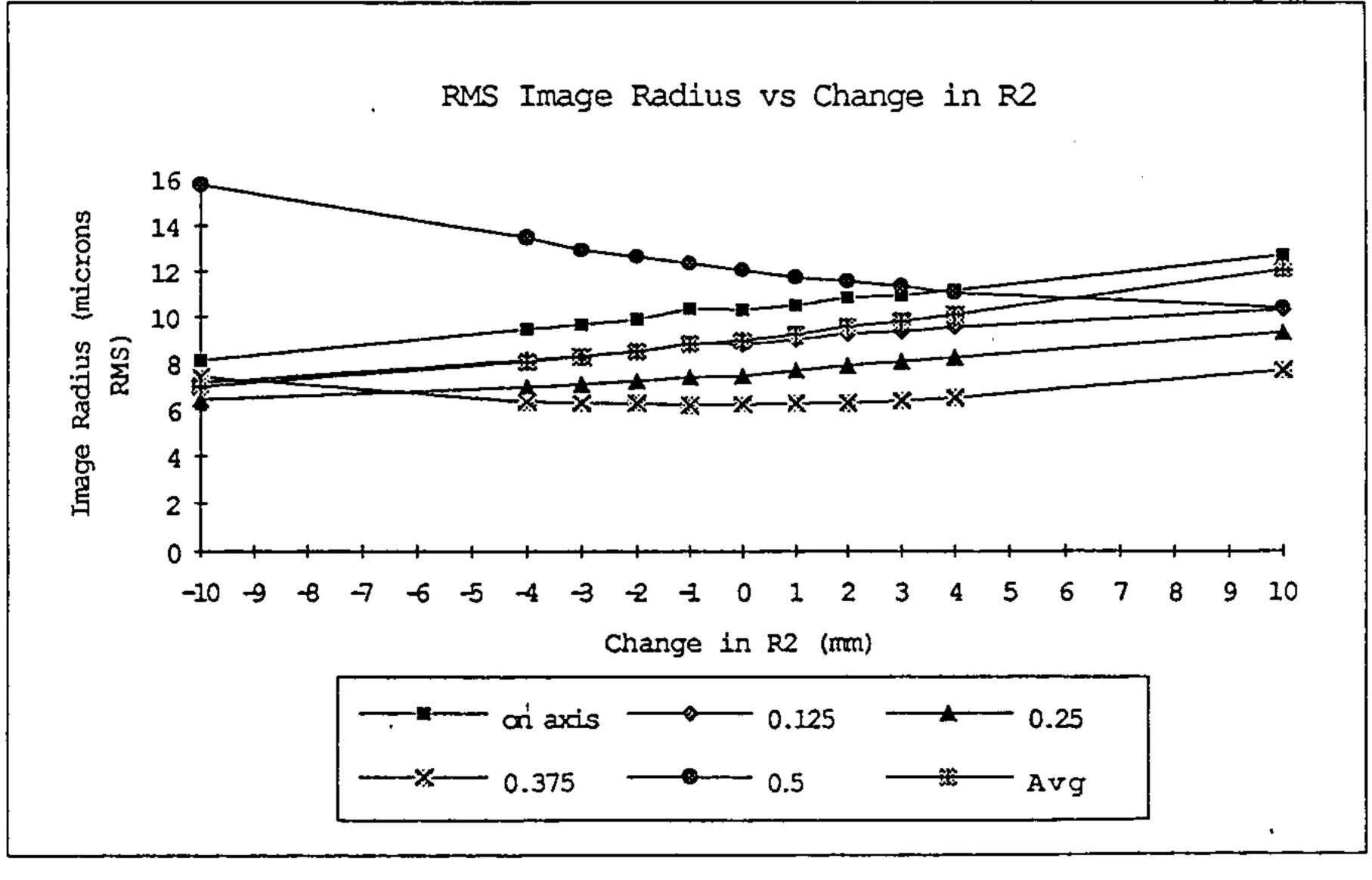
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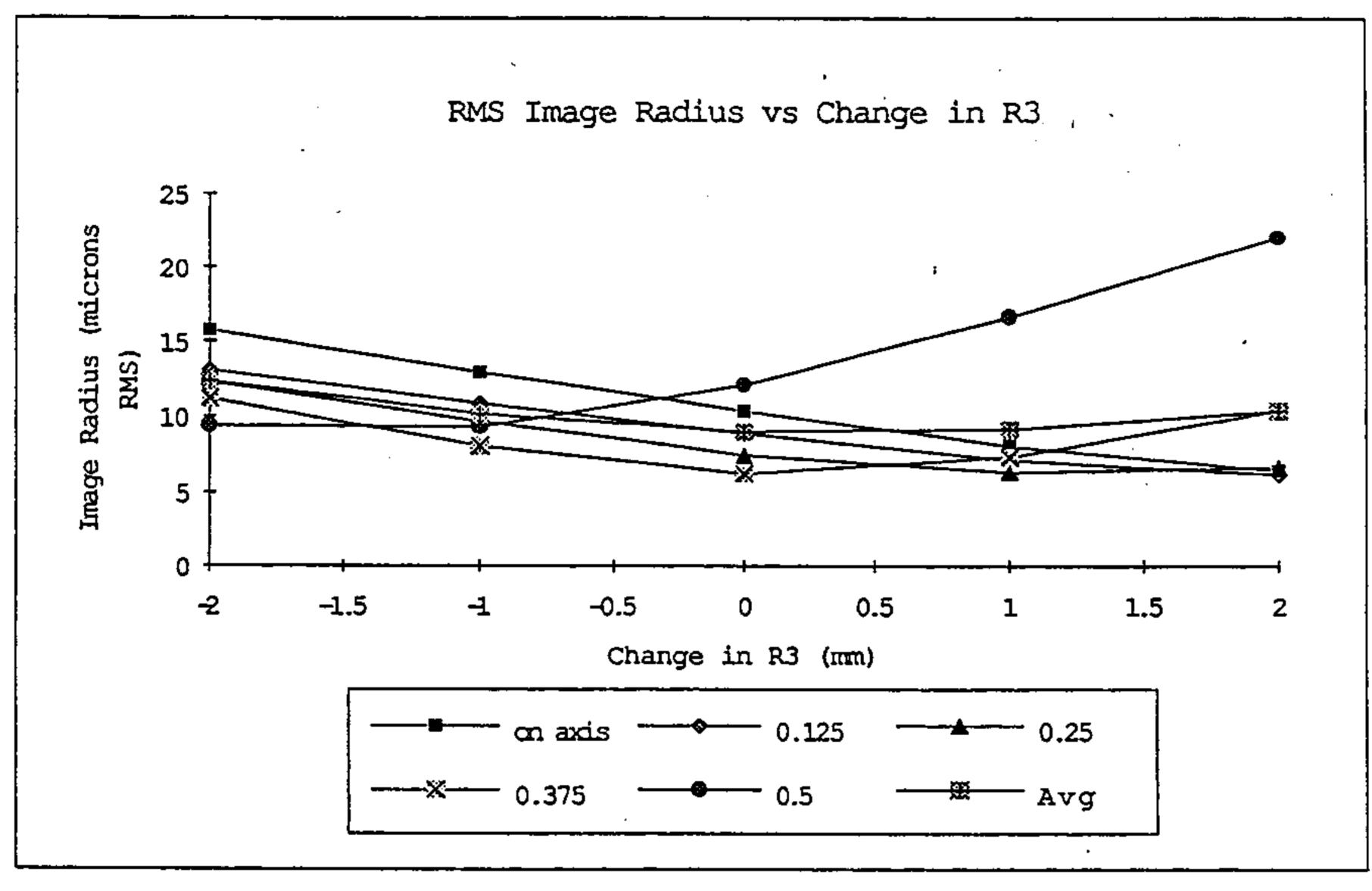
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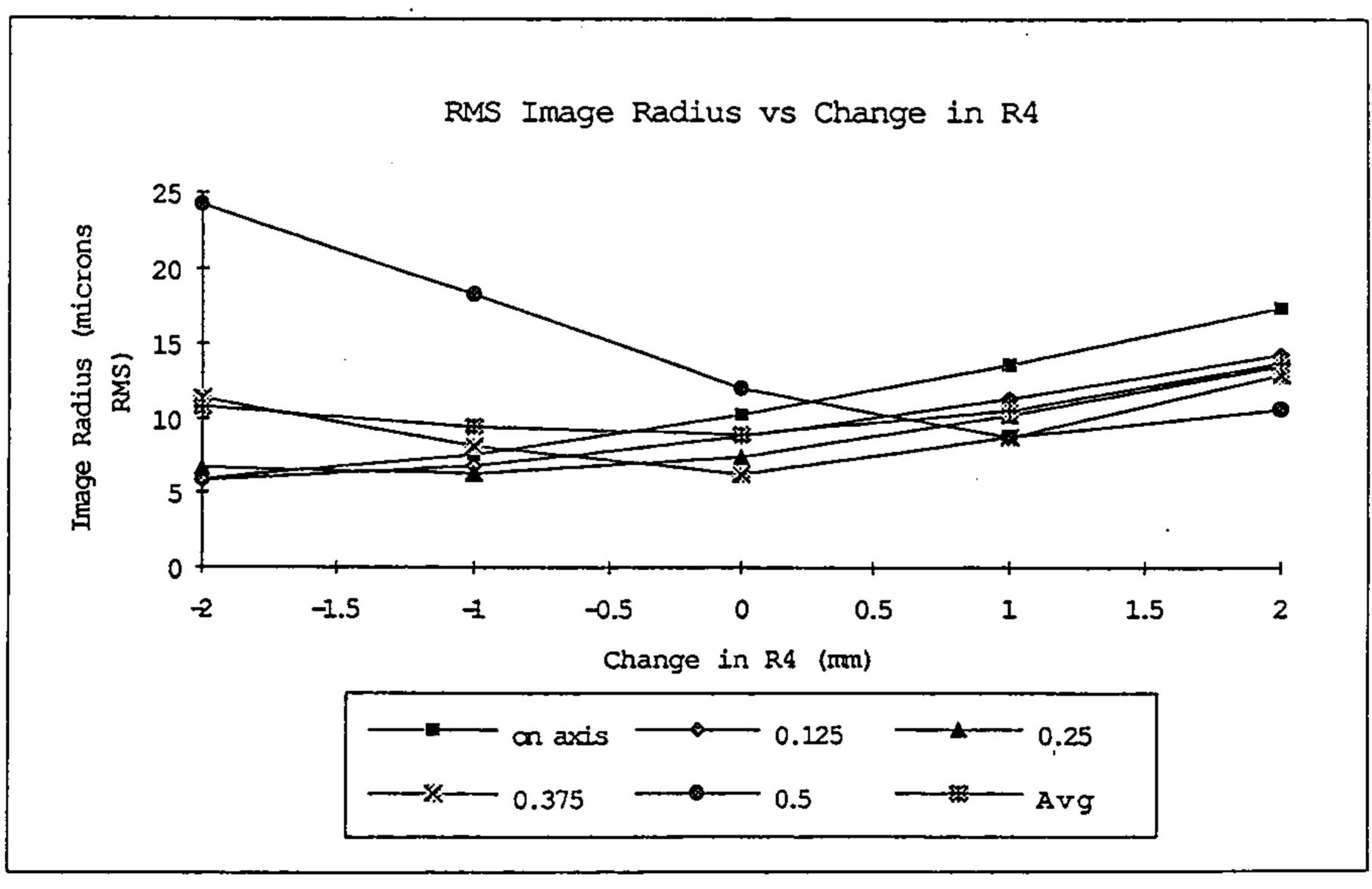


FIGURE 6